

PTEP

Progress of Theoretical and Experimental Physics

Review of Particle Physics

R.L. Workman *et al.* (Particle Data Group), *Prog. Theor. Exp. Phys.* 2022, 083C01 (2022)

PDG
particle data group



The Physical Society of Japan

OXFORD
UNIVERSITY PRESS

REVIEW OF PARTICLE PHYSICS*

Particle Data Group

Abstract

The *Review* summarizes much of particle physics and cosmology. Using data from previous editions, plus 2,143 new measurements from 709 papers, we list, evaluate, and average measured properties of gauge bosons and the recently discovered Higgs boson, leptons, quarks, mesons, and baryons. We summarize searches for hypothetical particles such as supersymmetric particles, heavy bosons, axions, dark photons, etc. Particle properties and search limits are listed in Summary Tables. We give numerous tables, figures, formulae, and reviews of topics such as Higgs Boson Physics, Supersymmetry, Grand Unified Theories, Neutrino Mixing, Dark Energy, Dark Matter, Cosmology, Particle Detectors, Colliders, Probability and Statistics. Among the 120 reviews are many that are new or heavily revised, including a new review on Machine Learning, and one on Spectroscopy of Light Meson Resonances.

The *Review* is divided into two volumes. Volume 1 includes the Summary Tables and 97 review articles. Volume 2 consists of the Particle Listings and contains also 23 reviews that address specific aspects of the data presented in the Listings.

The complete *Review* (both volumes) is published online on the website of the Particle Data Group (pdg.lbl.gov) and in a journal. Volume 1 is available in print as the *PDG Book*. A *Particle Physics Booklet* with the Summary Tables and essential tables, figures, and equations from selected review articles is available in print, as a web version optimized for use on phones, and as an Android app.

The 2022 edition of the *Review of Particle Physics* should be cited as:
R.L. Workman *et al.* (Particle Data Group), *Prog. Theor. Exp. Phys.* **2022**, 083C01 (2022)

DOI: 10.1093/ptep/ptac097

For the online version see pdg.lbl.gov:



© 2022

Except where otherwise noted, content of this work (the *Review of Particle Physics*) is licensed under a Creative Commons Attribution-NonCommercial 4.0 International (CC BY-NC 4.0) license.

*The publication of the *Review of Particle Physics* is supported by the Director, Office of Science, Office of High Energy Physics of the U.S. Department of Energy under Contract No. DE-AC02-05CH11231; by an implementing arrangement between the governments of Japan (MEXT: Ministry of Education, Culture, Sports, Science and Technology) and the United States (DOE) on cooperative research and development; by the Italian National Institute of Nuclear Physics (INFN); by the Physical Society of Japan (JPS); and by the European Laboratory for Particle Physics (CERN). Individual collaborators receive support for their PDG activities from their respective institutes or funding agencies.

Particle Data Group

R.L. Workman,¹ V.D. Burkert,² V. Crede,³ E. Klempt,⁴ U. Thoma,⁴ L. Tiator,⁵ K. Agashe,⁶ G. Aielli,⁷ B.C. Allanach,⁸ C. Amsler,⁹ M. Antonelli,¹⁰ E.C. Aschenauer,¹¹ D.M. Asner,¹¹ H. Baer,¹² Sw. Banerjee,¹³ R.M. Barnett,¹⁴ L. Baudis,¹⁵ C.W. Bauer,¹⁴ J.J. Beatty,¹⁶ V.I. Belousov,¹⁷ J. Beringer,¹⁴ A. Bettini*,¹⁸ O. Biebel,¹⁹ K.M. Black,²⁰ E. Blucher,²¹ R. Bonventre,¹⁴ V.V. Bryzgalov,¹⁷ O. Buchmuller,²² M.A. Bychkov,²³ R.N. Cahn,¹⁴ M. Carena,^{24,21,25} A. Ceccucci,²⁶ A. Cerri,²⁷ R. Sekhar Chivukula,²⁸ G. Cowan,²⁹ K. Cranmer,³⁰ O. Cremonesi,³¹ G. D'Ambrosio,³² T. Damour,³³ D. de Florian,³⁴ A. de Gouvêa,³⁵ T. DeGrand,³⁶ P. de Jong,³⁷ S. Demers,³⁸ B.A. Dobrescu,²⁴ M. D'Onofrio,³⁹ M. Doser,²⁶ H.K. Dreiner,⁴⁰ P. Eerola,⁴¹ U. Egede,⁴² S. Eidelman†,^{43,44} A.X. El-Khadra,⁴⁵ J. Ellis,^{46,26} S. C. Eno,⁶ J. Erler,⁵ V.V. Ezhela,¹⁷ W. Fetscher,⁴⁷ B.D. Fields,^{48,45} A. Freitas,⁴⁹ H. Gallagher,⁵⁰ Y. Gershtein,⁵¹ T. Gherghetta,⁵² M.C. Gonzalez-Garcia,^{53,54,55} M. Goodman,⁵⁶ C. Grab,⁴⁷ A.V. Gritsan,⁵⁷ C. Grojean,^{58,59} D.E. Groom,¹⁴ M. Grünewald,⁶⁰ A. Gurtu,^{26,61} T. Gutsche,⁶² H.E. Haber,⁶³ Matthieu Hamel,⁶⁴ C. Hanhart,⁶⁵ S. Hashimoto,⁶⁶ Y. Hayato,^{67,68} A. Hebecker,⁶⁹ S. Heinemeyer,⁷⁰ J. J. Hernández-Rey‡,⁷¹ K. Hikasa,^{72,73,74} J. Hisano,⁷⁵ A. Höcker,²⁶ J. Holder,^{76,77} L. Hsu,²⁴ J. Huston,⁷⁸ T. Hyodo,⁷⁹ Al. Ianni,⁸⁰ M. Kado,^{81,82,83} M. Karliner,⁸⁴ U.F. Katz,⁸⁵ M. Kenzie,⁸⁶ V.A. Khoze,⁸⁷ S.R. Klein,^{88,89} F. Krauss,⁸⁷ M. Kreps,⁸⁶ P. Krizan,^{90,91} B. Krusche†,⁹² Y. Kwon,⁹³ O. Lahav,⁹⁴ J. Laiho,⁹⁵ L.P. Lellouch,⁹⁶ J. Lesgourgues,⁹⁷ A.R. Liddle,⁹⁸ Z. Ligeti,¹⁴ C.-J. Lin,¹⁴ C. Lippmann,⁹⁹ T.M. Liss,¹⁰⁰ L. Littenberg,¹¹ C. Lourenço,²⁶ K.S. Lugovsky,^{14,17} S.B. Lugovsky,¹⁷ A. Lusiani,^{101,102} Y. Makida,⁶⁶ F. Maltoni,^{103,104} T. Mannel,¹⁰⁵ A.V. Manohar,²⁸ W.J. Marciano,¹¹ A. Masoni,¹⁰⁶ J. Matthews,¹⁰⁷ U.-G. Meißner,^{4,65} I.-A. Melzer-Pellmann,⁵⁸ M. Mikhasenko,¹⁹ D.J. Miller,¹⁰⁸ D. Milstead,¹⁰⁹ R.E. Mitchell,¹¹⁰ K. Mönig,¹¹¹ P. Molaro,^{112,113} F. Moortgat,^{26,114} M. Moskvic,²⁶ K. Nakamura,^{66,68} M. Narain,¹¹⁵ P. Nason,^{31,116} S. Navas‡,¹¹⁷ A. Nelles,^{111,85} M. Neubert,¹¹⁸ P. Nevski§,¹¹ Y. Nir,¹¹⁹ K.A. Olive,⁵² C. Patrignani,¹²⁰ J.A. Peacock,¹²¹ V.A. Petrov,¹⁷ E. Pianori,¹⁴ A. Pich,⁷¹ A. Piepke,¹²² F. Pietropaolo,^{26,123} A. Pomarol,^{124,125} S. Pordes,²⁴ S. Profumo,⁶³ A. Quadt,¹²⁶ K. Rabbertz,¹²⁷ J. Rademacker,¹²⁸ G. Raffelt,¹²⁹ M. Ramsey-Musolf,^{130,131,132} B.N. Ratcliff,¹³³ P. Richardson,⁸⁷ A. Ringwald,⁵⁸ D.J. Robinson,¹⁴ S. Roesler,²⁶ S. Rolli,¹³⁴ A. Romaniouk,^{135,136} L. J. Rosenberg,¹³⁷ J.L. Rosner,²¹ G. Rybka,¹³⁷ M.G. Ryskin,¹³⁸ R.A. Rytin,¹⁷ Y. Sakai,⁶⁶ S. Sarkar,¹³⁹ F. Sauli§,²⁶ O. Schneider,¹⁴⁰ S. Schönert,¹⁴¹ K. Scholberg,¹⁴² A.J. Schwartz,¹⁴³ J. Schwiening,⁹⁹ D. Scott,¹⁴⁴ F. Sefkow,⁵⁸ U. Seljak,^{89,14} V. Sharma,²⁸ S.R. Sharpe,¹³⁷ V. Shiltsev¶,²⁴ G. Signorelli,¹⁰² M. Silari,²⁶ F. Simon,¹²⁹ T. Sjöstrand,¹⁴⁵ P. Skands,⁴² T. Skwarnicki,⁹⁵ G.F. Smoot,^{146,89,14,147,148} A. Soffer,⁸⁴ M.S. Sozzi,¹⁴⁹ S. Spanier,¹⁵⁰ C. Spiering,¹¹¹ A. Stahl,¹⁵¹ S.L. Stone†,⁹⁵ Y. Sumino,⁷⁴ M.J. Syphers,^{152,24} F. Takahashi,⁷⁴ M. Tanabashi,^{153,75} J. Tanaka,¹⁵⁴ M. Taševský,¹⁵⁵ K. Terao,^{133,156} K. Terashi,¹⁵⁴ J. Terning,¹⁵⁷ R.S. Thorne,⁹⁴ M. Titov,¹⁵⁸ N.P. Tkachenko,¹⁷ D.R. Tovey,¹⁵⁹ K. Trabelsi,⁸³ P. Urquijo,¹⁶⁰ G. Valencia,⁴² R. Van de Water,²⁴ N. Varelas,¹⁶¹ G. Venanzoni,¹⁰² L. Verde,^{55,54} I. Vivarelli,²⁷ P. Vogel,¹⁶² W. Vogelsang,⁶² V. Vorobyev,^{43,44} S.P. Wakely,^{21,25} W. Walkowiak,¹⁰⁵ C.W. Walter,¹⁴² D. Wands,¹⁶³ D.H. Weinberg,¹⁶⁴ E.J. Weinberg,¹⁶⁵ N. Wermes,⁴⁰ M. White,^{89,14} L.R. Wiencke,¹⁶⁶ S. Willocq,¹³² C.G. Wohl,¹⁴ C.L. Woody,¹¹ W.-M. Yao,¹⁴ M. Yokoyama,^{167,68} R. Yoshida,⁵⁶ G. Zanderighi,¹⁶⁸ G.P. Zeller,²⁴ O.V. Zenin,^{17,169} R.-Y. Zhu,¹⁷⁰ Shi-Lin Zhu,¹⁷¹ F. Zimmermann,²⁶ P.A. Zyla¹⁴

Technical Associates: J. Anderson,¹⁴ T. Basaglia,²⁶ P. Schaffner,¹⁴ W. Zheng,¹⁷²

1. *George Washington University, Department of Physics, Washington, D.C., USA*
2. *Jefferson Lab, Newport News, VA, USA*
3. *Florida State University, Department of Physics, Tallahassee, FL, USA*
4. *Universität Bonn, Helmholtz-Institut für Strahlen- und Kernphysik, Bonn, Germany*
5. *Institut für Kernphysik, Johannes Gutenberg University, Mainz, Germany*
6. *University of Maryland, Department of Physics, College Park, MD, USA*
7. *Università degli Studi di Roma "Tor Vergata", Rome, Italy*
8. *Department of Applied Mathematics and Theoretical Physics, University of Cambridge, Cambridge, UK*
9. *Stefan Meyer Institute for Subatomic Physics, Austrian Academy of Sciences, Vienna, Austria*
10. *Lab. Nazionali di Frascati dell'INFN, Frascati, Italy*
11. *Brookhaven National Laboratory, Nuclear and Particle Physics Directorate, Upton, NY, USA*
12. *University of Oklahoma, Department of Physics and Astronomy, Norman, OK, USA*
13. *University of Louisville, Louisville, KY, USA*
14. *Physics Division, Lawrence Berkeley National Laboratory, Berkeley, CA, USA*
15. *Universität Zürich, Physik-Institut, Zürich, Switzerland*
16. *Ohio State University, Department of Physics, Columbus, OH, USA*

*Coordination activities supported directly by INFN.

†Deceased.

‡Support from Programa Estatal de Generación de Conocimiento MCIU, Spain and ERDF of the European Union (PGC2018-096663-B-C41 and C44).

§Retired.

¶Supported by Fermi Research Alliance, LLC under Contract No. DE-AC02-07CH11359 with DOE.

17. *Institute for High Energy Physics of the National Research Centre Kurchatov Institute, COMPAS Group, Protvino, Russia*
18. *INFN and Dipartimento di Fisica e Astronomia, Università di Padova, Padova, Italy*
19. *Ludwig-Maximilians-Universität, Fakultät für Physik, München, Germany*
20. *University of Wisconsin, Department of Physics, Madison, WI, USA*
21. *University of Chicago, Enrico Fermi Institute and Department of Physics, Chicago, IL, USA*
22. *Imperial College, High Energy Physics Group, Blackett Laboratory, London, UK*
23. *University of Virginia, Department of Physics, Charlottesville, VA, USA*
24. *Fermi National Accelerator Laboratory, Batavia, IL, USA*
25. *University of Chicago, Kavli Institute for Cosmological Physics, Chicago, IL, USA*
26. *CERN, European Organization for Nuclear Research, Genève, Switzerland*
27. *Department of Physics and Astronomy, University of Sussex, Brighton, UK*
28. *Department of Physics, University of California at San Diego, La Jolla, CA, USA*
29. *Department of Physics, Royal Holloway, University of London, London, UK*
30. *State University of New York, Institute for Theoretical Physics, Stony Brook, NY, USA*
31. *INFN Sezione di Milano-Bicocca, Piazza della Scienza, Milano, Italy*
32. *INFN Sezione di Napoli, Napoli, Italy*
33. *Institut des Hautes Etudes Scientifiques, Bures-sur-Yvette, France*
34. *UNSAM - Universidad Nacional de San Martin, International Center for Advanced Studies (ICAS) and Instituto de Ciencias Físicas (ICIFI), Buenos Aires, Argentina*
35. *Northwestern University, Department of Physics and Astronomy, Evanston, IL, USA*
36. *University of Colorado at Boulder, Department of Physics, Boulder, CO, USA*
37. *Nikhef and University of Amsterdam, Amsterdam, The Netherlands*
38. *Yale University, New Haven, CT, USA*
39. *University of Liverpool, Department of Physics, Liverpool, UK*
40. *Universität Bonn, Physikalisches Institut, Bonn, Germany*
41. *University of Helsinki, Department of Physics, Helsinki, Finland*
42. *Monash University, School of Physics and Astronomy, Melbourne, Australia*
43. *Budker Institute of Nuclear Physics SB RAS, Novosibirsk, Russia*
44. *Novosibirsk State University, Novosibirsk, Russia*
45. *University of Illinois, Department of Physics, Urbana, IL, USA*
46. *King's College London, Department of Physics, London, UK*
47. *ETH Zurich, Institute for Particle Physics and Astrophysics, Zurich, Switzerland*
48. *University of Illinois, Department of Astronomy, Urbana, IL, USA*
49. *University of Pittsburgh, Department of Physics and Astronomy, Pittsburgh, PA, USA*
50. *Tufts University, Department of Physics and Astronomy, Medford, MA, USA*
51. *Department of Physics and Astronomy, Rutgers University, NJ, USA*
52. *University of Minnesota, School of Physics and Astronomy, Minneapolis, MN, USA*
53. *CN Yang Institute for Theoretical Physics, Stony Brook University, Stony Brook, NY, USA*
54. *Institució Catalana de Recerca i Estudis Avancats, Barcelona, Spain*
55. *Instituto de ciencias del Cosmos (ICC), University of Barcelona, Barcelona, Spain*
56. *Argonne National Laboratory, Argonne, IL, USA*
57. *Johns Hopkins University, Baltimore, MD, USA*
58. *Deutsches Elektronen-Synchrotron DESY, Hamburg, Germany*
59. *Institut für Physik, Humboldt-Universität zu Berlin, Berlin, Germany*
60. *University College Dublin, School of Physics, Dublin, Ireland*
61. *TIFR, Mumbai, India*
62. *Universität Tübingen, Institut für Theoretische Physik, Tübingen, Germany*
63. *Santa Cruz Institute for Particle Physics, University of California, Santa Cruz, CA, USA*
64. *Université Paris-Saclay, CEA, LIST, F-91120 Palaiseau, France*
65. *Institut für Kernphysik and Institute for Advanced Simulation, Forschungszentrum Jülich, Jülich, Germany*
66. *KEK, High Energy Accelerator Research Organization, Tsukuba, Japan*
67. *Kamioka Observatory, ICRR, The University of Tokyo, Tokyo, Japan*
68. *The University of Tokyo, Kavli IPMU (WPI), The University of Tokyo Institutes for Advanced Study, Kashiwa, Japan*
69. *Heidelberg University, Institute for Theoretical Physics, Heidelberg, Germany*

70. *Instituto de Física Teórica (UAM/CSIC), Universidad Autónoma de Madrid, Madrid, Spain*
71. *IFIC — Instituto de Física Corpuscular, Universitat de València — C.S.I.C., Valencia, Spain*
72. *Division for Interdisciplinary Advanced Research and Education, Tohoku University, Sendai, Japan*
73. *Institute of Liberal Arts and Sciences, Tohoku University, Sendai, Japan*
74. *Tohoku University, Department of Physics, Sendai, Japan*
75. *Nagoya University, Kobayashi-Maskawa Institute, Nagoya, Japan*
76. *University of Delaware, Department of Physics and Astronomy, Newark, DE, USA*
77. *University of Delaware, Bartol Research Institute, Newark, DE, USA*
78. *Michigan State University, Dept. of Physics and Astronomy, East Lansing, MI, USA*
79. *Department of Physics, Tokyo Metropolitan University, Tokyo, Japan*
80. *INFN - Laboratori Nazionali del Gran Sasso, Assergi, Italy*
81. *Dipartimento di Fisica, Sapienza Università di Roma, Rome, Italy*
82. *INFN Sezione di Roma, Rome, Italy*
83. *IJCLab, CNRS/IN2P3, Université Paris-Saclay, Orsay, France*
84. *Department of Particle Physics, Tel-Aviv University, Tel Aviv, Israel*
85. *Friedrich-Alexander-Universität Erlangen-Nürnberg, Erlangen Centre for Astroparticle Physics, Erlangen, Germany*
86. *University of Warwick, Department of Physics, Coventry, UK*
87. *University of Durham, Institute for Particle Physics Phenomenology, Department of Physics, Durham, UK*
88. *Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, CA, USA*
89. *University of California, Department of Physics, Berkeley, CA, USA*
90. *Faculty of Mathematics and Physics, University of Ljubljana, Jadranska, Slovenia*
91. *Jozef Stefan Institute, Ljubljana, Slovenia*
92. *University of Basel, Institute of Physics, Basel, Switzerland*
93. *Yonsei University, Department of Physics, Seoul, Republic of Korea*
94. *University College London, Department of Physics and Astronomy, London, UK*
95. *Syracuse University, Department of Physics, Syracuse, NY, USA*
96. *Aix-Marseille Univ, Université de Toulon, CNRS, CPT, Marseille, France*
97. *Institute of Theoretical Particle Physics and Cosmology (TTK), RWTH, Aachen, Germany*
98. *Instituto de Astrofísica e Ciências do Espaço, Universidade de Lisboa, Lisbon, Portugal*
99. *GSI, Helmholtzzentrum für Schwerionenforschung, Darmstadt, Germany*
100. *The City College of New York, New York, NY, USA*
101. *Scuola Normale Superiore, Pisa, Italy*
102. *INFN Sezione di Pisa, Pisa, Italy*
103. *Université catholique de Louvain, Centre for Cosmology, Particle Physics and Phenomenology (CP3), Louvain-la-Neuve, Belgium*
104. *Università di Bologna and INFN, Dipartimento di Fisica e Astronomia, Bologna, Italy*
105. *Universität Siegen, Department für Physik, Siegen, Germany*
106. *INFN Sezione di Cagliari, Monserrato, Italy*
107. *Louisiana State University, Department of Physics and Astronomy, Baton Rouge, LA, USA*
108. *University of Glasgow, School of Physics and Astronomy, Glasgow, UK*
109. *Stockholms Universitet, AlbaNova University Centre, Fysikum, Stockholm, Sweden*
110. *Indiana University, Department of Physics, Bloomington, IN, USA*
111. *Deutsches Elektronen-Synchrotron DESY, Zeuthen, Germany*
112. *INAF-OATS, Trieste, Italy*
113. *Institute for Fundamental Physics of the Universe, Trieste, Italy*
114. *University of Ghent, Dept. of Physics and Astronomy, Ghent, Belgium*
115. *Brown University, Department of Physics, Providence, RI, USA*
116. *Dip. di Fisica "G. Occhialini", Università di Milano-Bicocca, Milano, Italy*
117. *Universidad de Granada, Dpto. de Física Teórica y del Cosmos & C.A.F.P.E., Granada, Spain*
118. *Johannes Gutenberg University, PRISMA Cluster of Excellence and Mainz Institute for Theoretical Physics, Mainz, Germany*
119. *Department of Particle Physics and Astrophysics, Weizmann Institute of Science, Rehovot, Israel*
120. *Università di Bologna and INFN, Dip. Scienze per la Qualità della Vita, Rimini, Italy*
121. *University of Edinburgh, Royal Observatory, Institute for Astronomy, Edinburgh, UK*
122. *University of Alabama, Department of Physics and Astronomy, Tuscaloosa, AL, USA*

123. INFN Sezione di Padova, Padua, Italy
124. Universitat Autònoma de Barcelona, Departament de Física, Barcelona, Spain
125. IFAE, Universitat Autònoma de Barcelona, Barcelona, Spain
126. Georg-August-Universität Göttingen, II. Physikalisches Institut, Göttingen, Germany
127. Karlsruhe Institute of Technology, Karlsruhe, Germany
128. University of Bristol, HH Wills Physics Laboratory, Bristol, UK
129. Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), München, Germany
130. Tsung-Dao Lee Institute, Shanghai Jiao Tong University, Shanghai, China
131. Shanghai Jiao Tong University, Shanghai, China
132. University of Massachusetts, Department of Physics, Amherst, MA, USA
133. SLAC National Accelerator Laboratory, Menlo Park, CA, USA
134. DOE, Washington, DC, USA
135. Institut für Astro- und Teilchenphysik, Leopold-Franzens-Universität, Innsbruck, Austria
136. National Research Nuclear University "MEPhI", Moscow, Russia
137. University of Washington, Department of Physics, Seattle, WA, USA
138. Petersburg Nuclear Physics Institute, Petersburg, Russia
139. University of Oxford, Rudolf Peierls Centre for Theoretical Physics, Oxford, UK
140. Institute of Physics, Ecole Polytechnique Fédérale de Lausanne (EPFL), Lausanne, Switzerland
141. Department of Physics, Technical University Munich, Munich, Germany
142. Duke University, Physics Department, Durham, NC, USA
143. University of Cincinnati, Department of Physics, Cincinnati, OH, USA
144. University of British Columbia, Department of Physics and Astronomy, Vancouver, BC, Canada
145. Lund University, Department of Astronomy and Theoretical Physics, Lund, Sweden
146. The Hong Kong University of Science and Technology, Kowloon, Hong Kong
147. Donostia International Physics Center (DIPC), Donostia-San Sebastián, Spain
148. Paris Centre for Cosmological Physics, APC (CNRS), Université de Paris, Paris, France
149. Pisa University, Pisa, Italy
150. University of Tennessee, Department of Physics and Astronomy, Knoxville, TN, USA
151. III. Physikalisches Institut, Physikzentrum, RWTH Aachen University, Aachen, Germany
152. Department of Physics, Northern Illinois University, DeKalb, IL, USA
153. Department of Physics, Nagoya University, Nagoya, Japan
154. International Center for Elementary Particle Physics (ICEPP), The University of Tokyo, Tokyo, Japan
155. Institute of Physics, Czech Academy of Sciences, Prague, Czech Republic
156. Stanford University, Stanford, CA, USA
157. Department of Physics, University of California, Davis, CA, USA
158. IRFU, CEA, Université Paris-Saclay, F-91191 Gif-sur-Yvette, France
159. University of Sheffield, Department of Physics and Astronomy, Sheffield, UK
160. University of Melbourne, School of Physics, Victoria, Australia
161. University of Illinois at Chicago, Chicago, IL, USA
162. California Institute of Technology, Kellogg Radiation Laboratory, Pasadena, CA, USA
163. University of Portsmouth, Institute of Cosmology and Gravitation, Portsmouth, UK
164. Ohio State University, Department of Astronomy and CCAPP, Columbus, OH, USA
165. Columbia University, Department of Physics, New York, NY, USA
166. Dept. of Physics, Colorado School of Mines, Golden, CO, USA
167. The University of Tokyo, Department of Physics, Tokyo, Japan
168. Max-Planck-Institute of Physics, Munich, Germany
169. Moscow Institute of Physics and Technology, Dolgoprudny, Russia
170. California Institute of Technology, High Energy Physics, Pasadena, CA, USA
171. School of Physics, Peking University, Beijing, China
172. Institute of High Energy Physics, Beijing, China

58. τ Branching Fractions

Revised February 2022 by Sw. Banerjee (Louisville U.) and A. Lusiani (SNS, Pisa; INFN, Pisa).

58.1 τ Branching Fractions

The τ Listings contains 244 entries that correspond to either a τ partial decay fraction into a specific decay mode (branching fraction) or a ratio of two τ partial decay fractions (branching ratio). Experimental information provides values for 147 of these quantities, upper limits for 61 branching fractions to Lepton Family number, Lepton number, or Baryon number violating modes, and 36 additional upper limits for other modes. A total of 170 measurements of τ branching fraction and branching ratio measurements is used for a global fit that determines 129 quantities.

58.2 The constrained fit to τ branching fractions

The τ branching fractions fit uses the reported values, uncertainties and statistical correlations of the τ branching fractions and branching ratios measurements. Asymmetric uncertainties are symmetrized as $\sigma_{\text{symm}}^2 = (\sigma_+^2 + \sigma_-^2)/2$. Additionally, the most precise experimental inputs are treated according to how they depend on external parameters on the basis of their documentation [1]. The τ measurements may depend on parameters such as the τ pair production cross-section in e^+e^- annihilations at the $\Upsilon(4S)$ peak. In some cases, measurements reported in different papers by the same collaboration may depend on common parameters like the estimate of the integrated luminosity or of particle identification efficiencies. For all the significant detected dependencies, the τ measurements and their uncertainties are updated to account for the updated values of the external parameters. The dependencies on common systematic effects are also determined in size and sign, and all the common systematic dependencies of different measurements are used together with the published statistical and systematic uncertainties and correlations in order to compute a single all-inclusive variance and covariance matrix of the experimental inputs of the fit.

The fit procedure parameters correspond to τ quantities that are fit to the experimental measurements while respecting relations described by a series of constraint equations. All the experimental inputs and all the constraint equations are reported in the τ Listings section that follows this review. With respect to the 2016, 2017 and 2018 editions, the fit uses one more experimental measurement, published by the BaBar collaboration in 2018, on $\mathcal{B}(\tau \rightarrow K^- K^0 \nu_\tau)$ [2]. If only a few measurements are correlated, the correlation coefficients are listed in the footnote for each measurement (see for example $\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K^0 \nu_\tau \text{ ("1-prong")})/\Gamma_{\text{total}}$). If a large number of measurements are correlated, then the full correlation matrix is listed in the footnote to the measurement that first appears in the τ Listings. Footnotes to the other measurements refer to the first one. For example, the large correlation matrices for the branching fraction or ratio measurements contained in Refs. [3] [4] are listed in Footnotes to the $\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ and $\Gamma(h^- \nu_\tau)/\Gamma_{\text{total}}$ measurements respectively. The constraints between the τ branching fractions and ratios include coefficients that correspond to physical quantities, like for instance the branching fractions of the η and ω mesons. All quantities are taken from the 2018 edition of the Review of Particle Physics. Their uncertainties are neglected in the fit.

We obtain the branching fraction of $\tau \rightarrow a_1^- (\rightarrow \pi^- \gamma) \nu_\tau$ using the ALEPH estimate for $\mathcal{B}(a_1^- \rightarrow \pi^- \gamma)$ [3], which uses the measurement of $\Gamma(a_1^- \rightarrow \pi^- \gamma)$ [5]. In the fit, we assume that $\mathcal{B}(\tau^- \rightarrow a_1^- \nu_\tau)$ is equal to $\mathcal{B}(\tau \rightarrow \pi^- \pi^- \pi^+ \nu_\tau \text{ (ex. } K^0, \omega)) + \mathcal{B}(\tau \rightarrow \pi^- 2\pi^0 \nu_\tau \text{ (ex. } K^0))$, neglecting the observed but negligible branching fractions to other modes, including $\mathcal{B}(a_1^- \rightarrow \pi^- \gamma)$.

In some cases, constraints describe approximate relations that nevertheless hold within the present experimental precision. For instance, the constraint $\mathcal{B}(\tau \rightarrow K^- K^- K^+ \nu_\tau) = \mathcal{B}(\tau \rightarrow K^- \phi \nu_\tau) \times \mathcal{B}(\phi \rightarrow K^+ K^-)$ is justified within the current experimental evidence.

In the fit, scale factors are applied to the published uncertainties of measurements only if significant inconsistency between different measurements remain after accounting for all relevant uncertainties and correlations. After examining the data and the fit pulls,

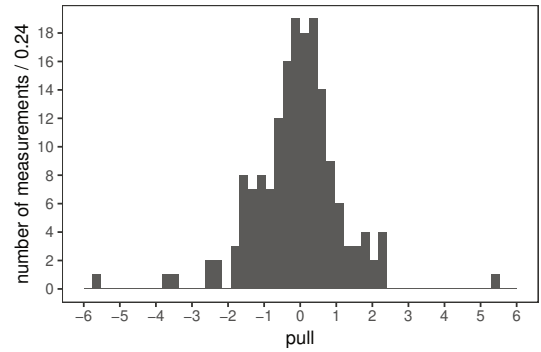


Figure 58.1: Pulls of individual measurements against the respective fitted quantity. No scale factor is used.

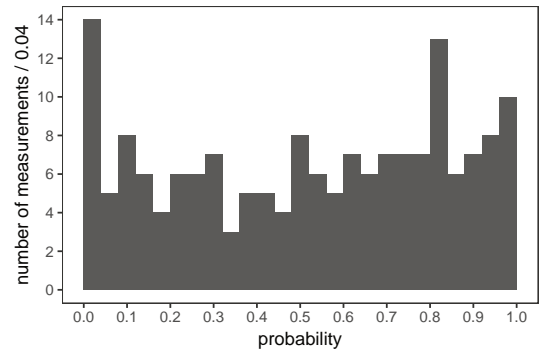


Figure 58.2: Probability of individual measurement pulls against the respective fitted quantity. No scale factor is used.

it has been decided to apply just one scale factor of 5.4 on the measurements of $\mathcal{B}(\tau \rightarrow K^- K^- K^+ \nu_\tau)$. The scale factor has been computed and applied according to the standard PDG procedure. Without the scale factor applied, the χ^2 probability of the fit is about 2%. On a per-measurement basis, the pull distribution in figure 58.1 indicates that just a few measurements have more than 3σ pulls. (The uncertainties to obtain the pulls are computed using the measurements variance matrix and the variance matrix of the result, accounting for the fact that the variance matrix of the result is obtained from the measurement variance with the fit.) The pull probability distribution in figure 58.2 is reasonably flat. With many measurements some entries on the tails of the normal distribution must be expected. There are 170 pulls, one per measurement. They are partially correlated, and the effective number of independent pulls is equal to the number of degrees of freedom of the fit, 125. Only the $\tau \rightarrow K^- K^- K^+ \nu_\tau$ decay mode has a pull that is inconsistent at the level of more than 3σ even if considered as the largest pull in a set of 125. This confirms the choice of adopting just that one scale factor.

After scaling the error, the constrained fit has a χ^2 of 135 for 125 degrees of freedom, corresponding to a χ^2 probability of 26%. We use 170 measurements and 84 constraints on the branching fractions and ratios to determine 129 quantities, consisting of 112 branching fractions and 17 branching ratios. A total of 85 quantities have at least one measurement in the fit. The constraints include the unitarity constraint on the sum of all the exclusive τ decay modes, $\mathcal{B}_{\text{all}} = 1$. If the unitarity constraint is released, the fit result for \mathcal{B}_{all} is consistent with unitarity with $1 - \mathcal{B}_{\text{all}} = (0.00 \pm 0.10)\%$.

For the convenience of summarizing the fit results, we list in the following the values and uncertainties for a set of 46 “basis” decay modes, from which all remaining branching fractions and ratios can be obtained using the constraints. The basis decay modes are not intended to sum up to 1. Since some basis quantities represent multiple branching fractions that are related by constraint equations, they are properly weighted and the unitarity constraint

corresponds to a linear combination whose coefficients are listed in the following. The correlation matrix between the basis modes is reported in the τ Listings.

decay mode	fit result (%)	coefficient
$\mu^- \bar{\nu}_\mu \nu_\tau$	17.3937 ± 0.0384	1.0000
$e^- \bar{\nu}_e \nu_\tau$	17.8175 ± 0.0399	1.0000
$\pi^- \nu_\tau$	10.8164 ± 0.0512	1.0000
$K^- \nu_\tau$	0.6964 ± 0.0096	1.0000
$\pi^- \pi^0 \nu_\tau$	25.4941 ± 0.0893	1.0000
$K^- \pi^0 \nu_\tau$	0.4328 ± 0.0148	1.0000
$\pi^- 2\pi^0 \nu_\tau$ (ex. K^0)	9.2595 ± 0.0964	1.0021
$K^- 2\pi^0 \nu_\tau$ (ex. K^0)	0.0647 ± 0.0218	1.0000
$\pi^- 3\pi^0 \nu_\tau$ (ex. K^0)	1.0429 ± 0.0707	1.0000
$K^- 3\pi^0 \nu_\tau$ (ex. K^0, η)	0.0478 ± 0.0212	1.0000
$h^- 4\pi^0 \nu_\tau$ (ex. K^0, η)	0.1118 ± 0.0391	1.0000
$\pi^- \bar{K}^0 \nu_\tau$	0.8384 ± 0.0138	1.0000
$K^- K^0 \nu_\tau$	0.1486 ± 0.0034	1.0000
$\pi^- \bar{K}^0 \pi^0 \nu_\tau$	0.3817 ± 0.0129	1.0000
$K^- \pi^0 K^0 \nu_\tau$	0.1500 ± 0.0070	1.0000
$\pi^- \bar{K}^0 2\pi^0 \nu_\tau$ (ex. K^0)	0.0263 ± 0.0226	1.0000
$\pi^- K_S^0 K_S^0 \nu_\tau$	0.0235 ± 0.0006	2.0000
$\pi^- K_S^0 K_L^0 \nu_\tau$	0.1081 ± 0.0241	1.0000
$\pi^- \pi^0 K_S^0 K_S^0 \nu_\tau$	0.0018 ± 0.0002	2.0000
$\pi^- \pi^0 K_S^0 K_L^0 \nu_\tau$	0.0325 ± 0.0119	1.0000
$\bar{K}^0 h^- h^- h^+ \nu_\tau$	0.0247 ± 0.0199	1.0000
$\pi^- \pi^- \pi^+ \nu_\tau$ (ex. K^0, ω)	8.9868 ± 0.0513	1.0021
$\pi^- \pi^- \pi^+ \pi^0 \nu_\tau$ (ex. K^0, ω)	2.7404 ± 0.0710	1.0000
$h^- h^- h^+ 2\pi^0 \nu_\tau$ (ex. K^0, ω, η)	0.0981 ± 0.0356	1.0000
$\pi^- K^- K^+ \nu_\tau$	0.1435 ± 0.0027	1.0000
$\pi^- K^- K^+ \pi^0 \nu_\tau$	0.0061 ± 0.0018	1.0000
$\pi^- \pi^0 \eta \nu_\tau$	0.1389 ± 0.0072	1.0000
$K^- \eta \nu_\tau$	0.0155 ± 0.0008	1.0000
$K^- \pi^0 \eta \nu_\tau$	0.0048 ± 0.0012	1.0000
$\pi^- \bar{K}^0 \eta \nu_\tau$	0.0094 ± 0.0015	1.0000
$\pi^- \pi^+ \pi^- \eta \nu_\tau$ (ex. K^0)	0.0220 ± 0.0013	1.0000
$K^- \omega \nu_\tau$	0.0410 ± 0.0092	1.0000
$h^- \pi^0 \omega \nu_\tau$	0.4085 ± 0.0419	1.0000
$K^- \phi \nu_\tau$	0.0044 ± 0.0016	0.8320
$\pi^- \omega \nu_\tau$	1.9494 ± 0.0645	1.0000
$K^- \pi^- \pi^+ \nu_\tau$ (ex. K^0, ω)	0.2927 ± 0.0068	1.0000
$K^- \pi^- \pi^+ \pi^0 \nu_\tau$ (ex. K^0, ω, η)	0.0394 ± 0.0142	1.0000
$\pi^- 2\pi^0 \omega \nu_\tau$ (ex. K^0)	0.0072 ± 0.0016	1.0000
$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. K^0, η, ω, f_1)	0.0014 ± 0.0027	1.0000
$3\pi^- 2\pi^+ \nu_\tau$ (ex. K^0, ω, f_1)	0.0775 ± 0.0030	1.0000
$K^- 2\pi^- 2\pi^+ \nu_\tau$ (ex. K^0)	0.0001 ± 0.0001	1.0000
$2\pi^- \pi^+ \omega \nu_\tau$ (ex. K^0)	0.0084 ± 0.0006	1.0000
$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. K^0, η, ω, f_1)	0.0038 ± 0.0009	1.0000
$K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. K^0)	0.0001 ± 0.0001	1.0000
$\pi^- f_1 \nu_\tau$ ($f_1 \rightarrow 2\pi^- 2\pi^+$)	0.0052 ± 0.0004	1.0000
$\pi^- 2\pi^0 \eta \nu_\tau$	0.0195 ± 0.0038	1.0000

In defining the fit constraints and in selecting the modes that sum up to one we made some assumptions and choices. We assume that some channels, like $\tau^- \rightarrow \pi^- K^+ \pi^- \geq 0\pi^0 \nu_\tau$ and $\tau^- \rightarrow \pi^+ K^- K^- \geq 0\pi^0 \nu_\tau$, have negligible branching fractions as expected from the Standard Model, even if the experimental limits for these branching fractions are not very stringent. The 95% confidence level upper limits are $\mathcal{B}(\tau^- \rightarrow \pi^- K^+ \pi^- \geq 0\pi^0 \nu_\tau) < 0.25\%$ and $\mathcal{B}(\tau^- \rightarrow \pi^+ K^- K^- \geq 0\pi^0 \nu_\tau) < 0.09\%$, values not so different from measured branching fractions for allowed 3-prong modes containing charged kaons. For decays to final states containing one neutral kaon we assume that the branching fraction with the K_L^0 are the same as the corresponding one with a K_S^0 . On decays with two neutral kaons we assume that the branching fractions with $K_L^0 K_L^0$ are the same as the ones with $K_S^0 K_S^0$.

58.3 BaBar and Belle measure on average lower branching fractions and ratios.

We compare the BaBar and Belle measurements with the results of a fit where all their measurements have been excluded. We find that that BaBar and Belle measure on average lower τ branching fractions and ratios than the other experiments. Figures 58.3 and

58.4 show histograms of the 28 normalized differences between the B -factory measurements and the respective non- B -factory fit results. The normalization is the uncertainty on the difference. The average normalized difference between the two sets of measurements is -0.8σ (-0.7σ for the 16 Belle measurements and -0.8σ for the 12 BaBar measurements).

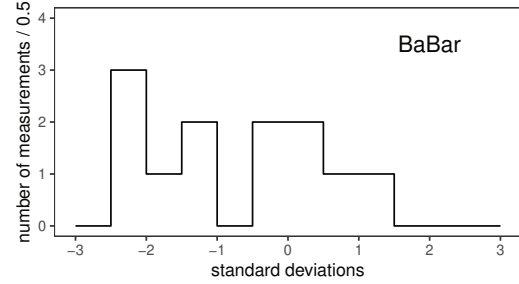


Figure 58.3: Distribution of the normalized difference between 12 measurements of branching fractions and ratios published by the BaBar collaboration and the respective averages computed using only non- B -factory measurements.

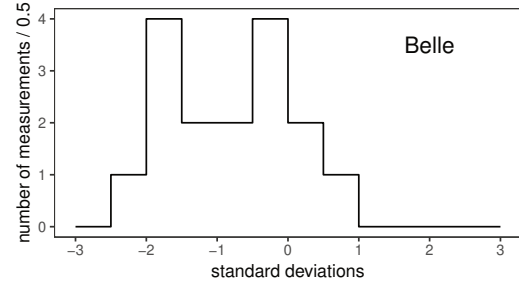


Figure 58.4: Distribution of the normalized difference between 16 measurements of branching fractions and ratios published by the Belle collaboration and the respective averages computed using only non- B -factory measurements.

58.4 Overconsistency of Leptonic Branching Fraction Measurements.

As observed in the previous editions of this review, measurements of the leptonic branching fractions are more consistent with each other than expected from the quoted errors on the individual measurements. The χ^2 is 0.34 for \mathcal{B}_e and 0.08 for \mathcal{B}_μ . Assuming normal errors, the probability of a smaller χ^2 is 1.3% for \mathcal{B}_e and 0.08% for \mathcal{B}_μ .

58.5 Technical implementation of the fit

The fit computes a set of quantities denoted with q_i by minimizing a χ^2 while respecting a series of equality constraints on the q_i . The χ^2 is computed using the measurements m_i and their covariance matrix E_{ij} as $\chi^2 = (m_i - A_{ik} q_k)^t E_{ij}^{-1} (m_j - A_{jl} q_l)$, where the model matrix A_{ij} is used to get the vector of the predicted measurements m'_i from the vector of the fit parameters q_j as $m'_i = A_{ij} q_j$. In this particular implementation the measurements are grouped by the quantity that they measure, and all quantities with at least one measurement correspond to a fit parameter. Therefore, the matrix A_{ij} has one row per measurement m_i and one column per fitted quantity q_j , with unity coefficients for the rows and column that identify a measurement m_i of the quantity q_j , respectively. The constraints are equations involving the fit parameters. The fit does not impose limitations on the functional form of the constraints. In summary, the fit requires:

$$\min [\chi^2(q_k)] = \min [(m_i - A_{ik} q_k)^t E_{ij}^{-1} (m_j - A_{jl} q_l)] \quad (58.1)$$

$$\text{subjected to } f_r(q_s) - c_r = 0 \quad (58.2)$$

where the left term of Eq. 58.2 defines the constraint expressions. Using the method of Lagrange multipliers, a set of equations is

obtained by taking the derivatives with respect to the fitted quantities q_k and the Lagrange multipliers λ_r of the sum of the χ^2 and the constraint expressions multiplied by the Lagrange multipliers λ_r , one for each constraint:

$$\begin{aligned} \min \left[(A_{ik}q_k - m_i)^t E_{ij}^{-1} (A_{jl}q_l - m_j) + 2\lambda_r (f_r(q_s) - c_r) \right] &= \\ &= \min \left[\tilde{\chi}^2(q_k, \lambda_r) \right] , \\ (\partial/\partial q_k, \partial/\partial \lambda_r) \left[\tilde{\chi}^2(q_k, \lambda_r) \right] &= 0 . \end{aligned} \quad (58.3)$$

Eq. 58.3 defines a set of equations for the vector of the unknowns (q_k, λ_r) , some of which may be non-linear, in case of non-linear constraints. An iterative minimization procedure approximates at each step the non-linear constraint expressions by their first order Taylor expansion around the current values of the fitted quantities, \bar{q}_s :

$$f_r(q_s) - c_r = f_r(\bar{q}_s) + \left. \frac{\partial f_r(q_s)}{\partial q_s} \right|_{\bar{q}_s} (q_s - \bar{q}_s) - c_r ,$$

which can be written as

$$B_{rs}q_s - c'_r ,$$

where c'_r are the resulting constant known terms, independent of q_s at first order. After linearization, the differentiation by q_k and λ_r is trivial and leads to a set of linear equations

$$A_{ki}^t E_{ij}^{-1} A_{jl}q_l + B_{kr}^t \lambda_r = A_{ki}^t E_{ij}^{-1} m_j , \quad (58.4)$$

$$B_{rs}q_s = c'_r , \quad (58.5)$$

which can be expressed as

$$F_{ij}u_j = v_i , \quad (58.6)$$

where $u_j = (q_k, \lambda_r)$ and v_i is the vector of the known constant terms running over the index k and then r in the right terms of Eq. 58.4 and Eq. 58.5, respectively. Solving the equation set in Eq. 58.6 by matrix inversion gives the fitted quantities and their variance and covariance matrix, using the measurements and their variance and covariance matrix. The fit procedure starts by computing the linear approximation of the non-linear constraint expressions around the quantities seed values. With an iterative procedure, the unknowns are updated at each step by solving the equations and the equations are then linearized around the updated values, until the variation of the fitted unknowns is reduced below a numerically small threshold.

References

- [1] D. Asner *et al.* (Heavy Flavor Averaging Group) (2010), [arXiv:1010.1589].
- [2] J. P. Lees *et al.* (BaBar), Phys. Rev. **D98**, 3, 032010 (2018), [arXiv:1806.10280].
- [3] S. Schael *et al.* (ALEPH), Phys. Rept. **421**, 191 (2005), [hep-ex/0506072].
- [4] J. Abdallah *et al.* (DELPHI), Eur. Phys. J. **C46**, 1 (2006), [hep-ex/0603044].
- [5] M. Zielinski *et al.*, Phys. Rev. Lett. **52**, 1195 (1984).