# Weyl Action of Two-Column Mixed-Symmetry Field and Its Factorization Around (A)dS Space

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ABSTRACT: We investigate the four-derivative free Weyl action for two-column mixed-symmetry field that makes use of maximal gauge symmetries. In flat space, the action can be uniquely determined from gauge and Weyl (trace shift) symmetry requirements. We show that there is a smooth and unique deformation of the flat action to (A)dS which keeps the same amount of gauge symmetries. This action admits a factorization into two distinct two-derivative actions having gauge parameters of different Young diagrams. Hence, this factorization pattern naturally extends that of the Weyl actions of symmetric higher spin fields to mixed-symmetry cases. The mass-deformation for these actions can be realized preserving one of the gauge symmetries. Although generically non-unitary, in special dimensions, unitarity is achieved selecting different mass deformations for dS and AdS. We consider particular examples of our construction such as New Massive Gravity in three dimensions, linearized bigravity in four dimensions and their arbitrary dimensional generalizations.

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#### 1 Introduction

In the realm of the standard field theories, more general fields than symmetric tensors are rarely used. It is partly because we are living in four-dimensional spacetime where any fundamental particle can be described by symmetric tensor(-spinor) fields. Hence, mixed-symmetry fields — fields whose indices have more general permutation symmetry than the fully symmetric ones — would be indispensable only in describing certain physics of higher dimensions, such as the physics of String Theory. Actually, the infinite tower of massive excitations in String Theory carry mixed-symmetry representations in general and their presence is crucial for the consistency of various stringy dualities. Moreover, mixed-symmetry fields might be useful, although not necessary, even in four dimensions, and their use may open a new avenue to unexplored land of physics.

The two-derivative Lagrangian for a massless mixed-symmetry field in flat spacetime is first given by Labastida [1–3] and then further investigated in [4–13]. Mixed-symmetry fields can be classified according to the symmetry property of the index permutations, hence can be associated with Young diagrams. The Lagrangian has in general a number of gauge symmetries, depending on the shape of the Young diagram associated with the field under consideration. One of the interesting and non-trivial properties of mixed-symmetry fields is that their flat space Lagrangian does not admit a smooth deformation towards the background with non-vanishing cosmological constant [14, 15]: around (A)dS background, there is no two-derivative mixed-symmetry Lagrangian respecting all the gauge symmetries available in flat space. Instead, one can choose to preserve only one gauge symmetry and the choice determines the mass squared term to a specific value in units of cosmological constant.

In this paper, we shall focus on the mixed-symmetry field associated with the Young diagrams having two columns. In a sense, the number of columns plays the role of 'spin' hence we are considering here only 'spin two' cases. Among generic two-columns cases, we will analyze the mixed-symmetry field corresponding to the simple hook Young diagram [16] in great details since it already contains all the essential features of the more general two-columns cases. The simple hook field is of particular interest because it appears in the first order formulation of Gravity through the spin connection. There have been various attempts to describe Gravity by the connection or related object (see for instance the recent discussion [17] and reference therein). We shall also discuss how these attempts can be understood from the perspective of the physics of mixed-symmetry field.

The main target of the current paper is the identification and the analysis of the higher-derivative action of mixed-symmetry field having Weyl transformation — that is, the trace shift — as its symmetry. In the more familiar case of symmetric spin-two field, the Weyl action is nothing but the linearization of the four-derivative Weyl gravity. The latter can be decomposed around (A)dS into the massless and partially-massless modes (see [18] for the related discussion). In the case of symmetric spin-s field, the Weyl action involves 2s derivatives which can be split into s different modes each of which is described by a specific two-derivative action with a certain gauge symmetry [19–21]: they are partially-massless spin s modes of depth 0 to s-1. In fact, these are all the known short representations (of

isometry group) containing the helicity s mode as the highest one.

This pattern suggests that the Weyl action of certain type of field — fully symmetric or mixed symmetry — could describe all the short representations associated with that field and each of these short representations can be realized as a two-derivative Lagrangian with a certain gauge symmetry (see [19] for some related discussions). In this paper, we examine this idea with the mixed-symmetry fields of two-column Young diagrams. The Weyl action of the field associated with a Young diagram having s columns ought to involve s derivatives. Hence, we expect it to describe s different short modes. The pattern of short representations involved might be non-trivial and the necessary analysis would be lengthy for generic Young diagrams. We reserve the general analysis for the future work and focus here only on the two-column cases where, irrespectively of the height of each column in the Young diagram, the Weyl action has four derivatives.

Let us provide more details on what has been done in this paper. We consider the fields having the symmetry of two-column Young diagram,

There are two short representations in (A)dS described by two distinct Lagrangians with different gauge symmetries: the gauge parameters have the index symmetry of Young diagram, either

$$p-1$$
 or  $p$  . (1.2)

In the flat limit, the two Lagrangians coincide and enjoy both of the gauge symmetries. Now considering the Weyl action of this field in flat space, we find that there is a unique four-derivative action invariant with respect to both gauge symmetries and a trace shift symmetry with the parameter,

$$\begin{array}{c|c}
 & q-1 \\
\hline
\end{array} \quad . \tag{1.3}$$

We show that this flat space Weyl action now admits a smooth deformation to the background with non-vanishing cosmological constant, as opposed to the two-derivative one. Moreover, this four-derivative action describes two short-representation modes corresponding to the gauge symmetries (1.2). In the case where the heights of two columns coincide, the Weyl action describes, on top of the usual massless field, new degrees of freedom (DoF) having a two-derivative gauge symmetry similar to the partially-massless spin two field.

Therefore, our analysis shows that the two-column cases precisely fit in the pattern discussed before.

Further, we study two-derivative massive deformations of the Weyl action. In (A)dS there are two distinct mass deformations, which preserve one or the other of the symmetries (1.2). We show that the massive action does not admit a smooth (A)dS deformation, similarly to the conventional two-derivative actions. Two different deformations have different spectra around (A)dS background. We observe that the sign of the coefficients in front of the free actions depend on the sign of the cosmological constant. In special dimension d = p + q + 1, one of the massive deformations is unitary around dS and non-unitary around AdS, while the other deformation follows the opposite pattern — unitary around AdS and non-unitary around dS. In other words, the unitarity requirement selects one of the massive deformations, and this choice is different for different signs of cosmological constant.

The plan of the paper is as follows. We focus first on the simple hook case in Section 2. After a detailed review of the two-derivative systems in Section 2.1 and 2.2, we construct and discuss the four-derivative Weyl action in Section 2.3. The result of the simple hook is generalized in Section 3 towards more general two-column cases. We provide separate discussions on the different height case (Section 3.1) and the equal height case (Section 3.2). Finally, we provide various discussions, in particular, related to New Massive Gravity [22–25], in Section 3.3.

# 2 Simple Hook

Let us first consider the simplest mixed-symmetry field  $\phi_{\mu\nu,\rho}$ , corresponding to the simple hook Young diagram,

$$\begin{array}{c|c} \mu & \rho \\ \nu & \end{array} \tag{2.1}$$

The shape of Young diagram dictates the symmetries of the field under index permutations. The symmetries are

$$\phi_{\mu\nu,\rho} = -\phi_{\nu\mu,\rho}, \qquad \phi_{\mu\nu,\rho} + \phi_{\nu\rho,\mu} + \phi_{\rho\mu,\nu} = 0.$$
 (2.2)

Notice that we work in the base where the anti-symmetry of indices in each column is manifest. One can equally work with the symmetric base, but for the construction of the Lagrangian we find the antisymmetric base more advantageous.

The Young diagram is fixing not only how fields transform under index permutations but also under Lorentz transformations. Hence, it defines a representation under Lorentz group SO(1, d-1) that the off-shell field carries.

#### 2.1 Einstein Action in Flat space

In flat spacetime, one can find a proper set of on-shell conditions — or, an action principle — that makes the field carry the same Young diagram representation but now under the little group SO(d-1) or SO(d-2), depending on whether the field is massive or massless<sup>1</sup>,

<sup>&</sup>lt;sup>1</sup>Only the representations of compact subgroup SO(d-2) of massless little group ISO(d-2) are relevant for our discussion in this paper.

respectively. When given Young diagram representation of Lorentz group does not exist for the little group, an interesting mechanism may emerge when the propagating modes carry certain Young diagram representations of little group which are different from the Young diagram representation of the Lorentz group that the off-shell field carries. For any mixed-symmetry Young diagram this happens in space-time dimensions lower than certain value. We shall return to this point later, but work for a while with an arbitrary d.

In [16], Curtright constructed the free action for the hook field  $\phi_{\mu\nu,\rho}$  which describes massless DoF carrying the hook Young diagram representation of the little group SO(d-2). Note that such Young diagram exists only when d is greater than four. By introducing the scalar product,

$$\langle A | B \rangle = \frac{1}{m! \, n!} \, \int d^d x \sqrt{|g|} \, A^{\mu_1 \cdots \mu_m, \nu_1 \cdots \nu_n}(x) \, B_{\mu_1 \cdots \mu_m, \nu_1 \cdots \nu_n}(x) \,, \tag{2.3}$$

the action for the hook field can be written as

$$S_{E}^{\text{flat}}[\phi] = \langle \phi \mid \mathcal{G} \phi \rangle. \tag{2.4}$$

Here the 'Einstein tensor'  $\mathcal{G} \phi$  is defined through the 'Ricci tensor'  $\mathcal{F} \phi$  as

$$(\mathcal{G}\,\phi)_{\mu\nu,\rho} = (\mathcal{F}\,\phi)_{\mu\nu,\rho} - \eta_{\rho[\mu}\,(\mathcal{F}\,\phi)_{\nu]\lambda,}^{\lambda}. \tag{2.5}$$

In this paper, the round/square brackets denote full symmetrization/anti-symmetrization with weight one, e.g.  $A_{[[\mu\nu]]} = A_{[\mu\nu]} = -A_{[\nu\mu]}$ . The 'Ricci tensor' itself is given by

$$(\mathcal{F}\phi)_{\mu\nu,\rho} = \partial^2 \phi_{\mu\nu,\rho} + 2 \,\partial_{[\mu} \,\partial^{\lambda} \phi_{\nu]\lambda,\rho} - \partial_{\rho} \,\partial^{\lambda} \phi_{\mu\nu,\lambda} - 2 \,\partial_{\rho} \,\partial_{[\mu} \phi_{\nu]\lambda,}{}^{\lambda}, \qquad (2.6)$$

which extends the Fierz-Pauli massless spin two equation to the hook field. By making use of the generalized Kronecker delta,

$$\delta^{\rho\lambda\gamma\delta}_{\mu\nu\alpha\beta} = 4! \, \delta^{\rho}_{[\mu} \, \delta^{\lambda}_{\nu} \, \delta^{\gamma}_{\alpha} \, \delta^{\delta}_{\beta]} \,, \tag{2.7}$$

the Einstein tensor can be written in a more compact form as

$$(\mathcal{G}\,\phi)_{\mu\nu,}{}^{\rho} = -\frac{1}{2}\,\delta^{\rho\lambda\gamma\delta}_{\mu\nu\alpha\beta}\,\partial^{\alpha}\,\partial_{\lambda}\,\phi_{\gamma\delta,}{}^{\beta}\,. \tag{2.8}$$

The action (2.4) has two gauge symmetries,

$$\delta_{\epsilon} \phi_{\mu\nu,\rho} = 2 \,\partial_{[\mu} \,\epsilon_{\nu]\rho}, \qquad \delta_{\theta} \,\phi_{\mu\nu,\rho} = \partial_{\rho} \,\theta_{\mu\nu} - \partial_{[\mu} \,\theta_{\nu]\rho}, \tag{2.9}$$

generated by symmetric and anti-symmetric gauge parameters:

$$\epsilon_{\mu\nu} = \epsilon_{\nu\mu} \,, \qquad \theta_{\mu\nu} = -\theta_{\nu\mu} \,.$$
 (2.10)

The above symmetry can be understood diagrammatically as

$$\delta_{\epsilon} \frac{\phi \phi}{\phi} = \frac{\epsilon \epsilon}{\partial}, \qquad \delta_{\theta} \frac{\phi \phi}{\phi} = \frac{\theta \partial}{\theta}. \tag{2.11}$$

This gauge symmetry is reducible admitting the gauge-for-gauge symmetry,

$$\epsilon_{\mu\nu}(\xi) = \partial_{(\mu} \, \xi_{\nu)} \,, \qquad \theta_{\mu\nu}(\xi) = \partial_{[\mu} \, \xi_{\nu]} \,,$$
 (2.12)

which reads in terms of Young diagram,

$$\begin{array}{c|c}
\hline
\epsilon & \epsilon
\end{array} = \begin{bmatrix}
\xi & \partial
\end{array}, \qquad \begin{bmatrix}
\theta \\
\theta
\end{bmatrix} = \begin{bmatrix}
\xi \\
\partial
\end{bmatrix}.$$
(2.13)

The gauge symmetries generated by  $\epsilon_{\mu\nu}(\xi)$  and  $\theta_{\mu\nu}(\xi)$  have the same form,

$$\frac{|\xi|}{|\partial|}, \qquad (2.14)$$

and they sum up to zero for appropriate choice of relative coefficient.

The structure of gauge symmetry tells us the number of DoF of this system. According to the covariant counting (see e.g [26]), we get

$$\mathcal{S}_{\mathrm{E}}$$
:  $\Box_{GL(d)} \ominus 2 \left( \Box_{GL(d)} \oplus \Box_{GL(d)} \right) \ominus 3 \Box_{GL(d)},$  (2.15)

where we subtract twice the DoF associated to gauge parameters as usual and put back three times of the DoF of gauge-for-gauge parameter. The above can be recast into the counting in the Hamiltonian analysis,

$$\square_{SO(d-1)} \ominus \left(\square_{SO(d-1)} \oplus \square_{SO(d-1)}\right) \oplus \square_{SO(d-1)}, \qquad (2.16)$$

where we subtract once the DoF for each of traceless gauge parameters and put back once that of gauge-for-gauge. This again can be rearranged into

which is the DoF of massive hook minus those of massless fields  $\square$ ,  $\square$  and  $\square$ . Hence, this gives the interpretation of obtaining massless hook from the massless limit of massive hook by eliminating other 'lower spin' components the latter involves. Eventually, we end up with the DoF of the simple hook of the little group SO(d-2). Hence, the DoF of the hook field can be conveniently counted, in flat space, through the number of components of the Young diagram representations of the little group.

### 2.2 Einstein Action in (A)dS

Let us move on to the (A)dS background,<sup>2</sup> and consider the analogous Lagrangian to the flat space one (2.4). One can think of the same expression for the Lagrangian as in (2.4) where all flat partial derivatives  $\partial_{\mu}$  are replaced by the (A)dS covariant derivatives  $\nabla_{\mu}$ , but this definition of Lagrangian has an ambiguity coming from the commutators of  $\nabla_{\mu}$ 's which give terms proportional to (A)dS curvature. The ambiguous term is hence a non-derivative

<sup>&</sup>lt;sup>2</sup>Explicit analysis making use of Stueckelberg fields can be found in [27].

mass-like term proportional to the cosmological constant, and it ought to be determined by the invariance of the action with respect to either of the gauge symmetries,

$$\delta_{\epsilon}^{\Lambda} \phi_{\mu\nu,\rho} = 2 \nabla_{[\mu} \epsilon_{\nu]\rho}, \qquad \delta_{\theta}^{\Lambda} \phi_{\mu\nu,\rho} = \nabla_{\rho} \theta_{\mu\nu} - \nabla_{[\mu} \theta_{\nu]\rho}. \tag{2.18}$$

By examining the gauge invariance with a possible mass-like term, one can realize that it is impossible to preserve both of the gauge symmetries. Depending on which symmetry we want to keep, the mass-like term is determined with a different mass-squared coefficient [14, 15].

Let us explictly determine the mass-term or equivalently the Einstein tensor  $\mathcal{G}_{\mu\nu,\rho}$  of the (A)dS Lagrangian defined in the same manner as in (2.4). To proceed, let us first fix our conventions:

$$\left[\nabla_{\mu}, \nabla_{\nu}\right] V_{\lambda}^{\rho} = R_{\mu\nu,\lambda}{}^{\sigma} V_{\sigma}^{\rho} - R_{\mu\nu,\sigma}{}^{\rho} V_{\lambda}^{\sigma}, \qquad R_{\mu\nu,\lambda}{}^{\rho} = \frac{2\Lambda}{(d-1)(d-2)} \left(g_{\mu\lambda}^{\Lambda} \delta_{\nu}^{\rho} - g_{\nu\lambda}^{\Lambda} \delta_{\mu}^{\rho}\right), \tag{2.19}$$

where  $\Lambda$  is the cosmological constant and  $g^{\Lambda}$  is the (A)dS metric.

By requiring the Lagrangian to be invariant under the transformation  $\delta_{\epsilon}^{\Lambda}$ , we determine the Eintein tensor as

$$(\mathcal{G}^{\Lambda}_{\square}\phi)_{\mu\nu,}{}^{\rho} = -\frac{1}{2} \,\delta^{\rho\lambda\gamma\delta}_{\mu\nu\alpha\beta} \,\nabla^{\alpha} \,\nabla_{\lambda} \,\phi_{\gamma\delta,}{}^{\beta} \,, \tag{2.20}$$

which has exactly the same form as the flat space one (2.8) except for the replacement of  $\partial_{\mu}$  by  $\nabla_{\mu}$ . Note that, as opposed to flat space case, the order of derivatives is important in the expression (2.20). This system admits a gauge-for-gauge symmetry

$$\epsilon_{\mu\nu} = \left(\nabla_{(\mu} \nabla_{\nu)} + \frac{2\Lambda}{(d-1)(d-2)} g^{\Lambda}_{\mu\nu}\right) \zeta, \qquad (2.21)$$

which corresponds to the partially-massless spin-two transformation.

On the other hand, for the invariance under the transformation  $\delta_{\theta}^{\Lambda}$ , the Einstein operator should acquire a mass-like term,

$$\mathcal{G}^{\Lambda}_{\exists} = \mathcal{G}^{\Lambda}_{\Box} - m^{2}_{\Lambda} \mathcal{I}, \tag{2.22}$$

where the mass-squared parameter is fixed by

$$m_{\Lambda}^2 = -\frac{4(d-3)}{(d-1)(d-2)}\Lambda,$$
 (2.23)

and the mass-term operator by

$$(\mathcal{I}\phi)_{\mu\nu,}{}^{\rho} = \frac{1}{2} \delta^{\rho\beta\gamma}_{\mu\nu\alpha} \phi_{\beta\gamma,}{}^{\alpha} = \phi_{\mu\nu,}{}^{\rho} - 2 \delta^{\rho}_{[\nu} \phi_{\mu]\alpha,}{}^{\alpha}. \tag{2.24}$$

Both of the actions  $\mathcal{G}^{\Lambda}_{\square}$  and  $\mathcal{G}^{\Lambda}_{\mathbb{H}}$  vanish identically for  $d \leq 3$ .

Notice that the mass-squared parameter is positive in AdS and negative in dS. In both of cases, the equation with higher mass-squared term is unitary —  $\mathcal{G}_{\mathbb{H}}^{\Lambda}$  for AdS and  $\mathcal{G}_{\mathbb{H}}^{\Lambda}$  for dS — whereas the other is non-unitary. See [27] for more details.

To recapitulate, the simple hook field in (A)dS space cannot have a two-derivative action respecting both of the gauge symmetries (2.18), but either the action  $\mathcal{S}_{E\square}^{\Lambda}$  with only the symmetric parameter gauge symmetry or the action  $\mathcal{S}_{E\square}^{\Lambda}$  with only the anti-symmetric parameter gauge symmetry:

$$S_{\text{Em}}^{\Lambda}[\phi] = \langle \phi \, | \, \mathcal{G}_{\text{m}}^{\Lambda} \, \phi \rangle \,, \qquad S_{\text{E}}^{\Lambda}[\phi] = \langle \phi \, | \, \mathcal{G}_{\text{H}}^{\Lambda} \, \phi \rangle \,. \tag{2.25}$$

Let us examine the DoF of the above systems. We first consider the action  $\mathcal{S}^{\Lambda}_{\text{Em}}$ . The DoF of the system can be counted in terms of the GL(d) Lorentz covariant tensors as

$$\mathcal{S}_{\mathrm{E}_{\square}}^{\Lambda}$$
:  $GL(d) \oplus 2 \square_{GL(d)} \oplus 3 \bullet$   $GL(d) \oplus \square_{GL(d)},$  (2.26)

which has a non-trivial pattern due to the mixture of first- and second-class constraints as well as gauge-for-gauge. Instead, in the Hamiltonian analysis, we find a simpler pattern,

$$SO(d-1) \ominus SO(d-1) \oplus \bullet, \qquad (2.27)$$

where we simply remove once the DoF of the traceless gauge parameter and put back that of gauge-for-gauge. We can further decompose these DoF as propagation on light-cone, ending up with

$$S_{\text{Em}}^{\Lambda}$$
:  $S_{O(d-2)} \oplus S_{O(d-2)}$ . (2.28)

Similarly, the number of DoF of the action  $\mathcal{S}_{\text{EH}}^{\Lambda}$  can be counted in terms of GL(d) covariant tensors as

$$\mathcal{S}_{\mathrm{EH}}^{\Lambda}$$
:  $\bigoplus_{GL(d)} \ominus 2 \bigoplus_{GL(d)} \ominus GL(d)$   $\ominus \bigoplus_{GL(d)} \oplus \bigoplus_{GL(d)} .$  (2.29)

The Hamiltonian analysis gives the same result, in terms of SO(d-1) branching:

$$SO(d-1) \ominus SO(d-1). \tag{2.30}$$

In terms of light-cone DoF, the equation (2.22) propagates

$$\mathcal{S}_{\mathrm{EH}}^{\Lambda}$$
:  $SO(d-2) \oplus SO(d-2)$ . (2.31)

Hence, compared to the flat space case, the hook action  $\mathcal{S}_{\text{E}\square}^{\Lambda}$  describes extra DoF corresponding to the light-cone propagation of a massless two-form field  $\square$ , while  $\mathcal{S}_{\text{E}\square}^{\Lambda}$  describes extra DoF of massless spin-two field  $\square$ . The general pattern of the decomposition of modes in flat limit for an arbitrary mixed-symmetry field has been conjectured in [28] and proved in [29, 30]. Let us note that in four dimensions, the hook mode identically vanishes hence the action  $\mathcal{S}_{\text{E}\square}^{\Lambda}$  describes only the massless spin two mode. In [17], it has been shown that this action can be obtained from the linearized Einstein-Cartan gravity in AdS<sub>4</sub> by integrating out the vielbein instead of the spin connection.

#### 2.3 Weyl Action

Even though there is no two-derivative action preserving both of the gauge symmetries of (2.18) in (A)dS, there may exist a higher-derivative action which is invariant with respect to both of the symmetries.

It turns out that the four-derivative action invariant under both symmetries has a simple form,

$$S_{\mathbf{W}}^{\Lambda}[\phi] = -\langle \phi \,|\, \mathcal{G}_{\square}^{\Lambda} \,\mathcal{I}^{-1} \,\mathcal{G}_{\mathbb{H}}^{\Lambda} \,\phi \rangle \,, \tag{2.32}$$

where the overall sign is chosen for the positive definite Euclidean action. This action makes use of both Einstein tensors as well as the inverse mass-term operator  $\mathcal{I}^{-1}$  given by

$$(\mathcal{I}^{-1}\phi)_{\mu\nu,}{}^{\rho} = \phi_{\mu\nu,}{}^{\rho} - \frac{2}{d-2} \delta^{\rho}_{[\nu} \phi_{\mu]\alpha,}{}^{\alpha}. \tag{2.33}$$

All the operators  $\mathcal{G}^{\Lambda}_{\square}$ ,  $\mathcal{G}^{\Lambda}_{\square}$  and  $\mathcal{I}$  are self-adjoint with respect to the scalar product (2.3):

$$\langle f \mid \mathcal{O} g \rangle = \langle \mathcal{O} f \mid g \rangle, \qquad \mathcal{O} = \mathcal{G}_{\square}^{\Lambda}, \ \mathcal{G}_{\mathbb{H}}^{\Lambda}, \ \mathcal{I}.$$
 (2.34)

Since the Einstein operators  $\mathcal{G}^{\Lambda}_{\square}$  and  $\mathcal{G}^{\Lambda}_{\mathbb{H}}$  differ by a mass-like term, it is easy to show that

$$S_{\mathbf{W}}^{\Lambda}[\phi] = -\langle \phi \mid \left( \mathcal{G}_{\mathbf{m}}^{\Lambda} \mathcal{I}^{-1} \mathcal{G}_{\mathbf{m}}^{\Lambda} - m_{\Lambda}^{2} \mathcal{G}_{\mathbf{m}}^{\Lambda} \right) \phi \rangle$$

$$= -\langle \phi \mid \left( \mathcal{G}_{\mathbf{H}}^{\Lambda} \mathcal{I}^{-1} \mathcal{G}_{\mathbf{H}}^{\Lambda} + m_{\Lambda}^{2} \mathcal{G}_{\mathbf{H}}^{\Lambda} \right) \phi \rangle. \tag{2.35}$$

In the first line, the action is manifestly invariant under the gauge symmetry  $\delta_{\epsilon}^{\Lambda}$  with the symmetric parameter, whereas the second line is manifestly invariant under the  $\delta_{\theta}^{\Lambda}$  one with anti-symmetric parameter.

In the flat limit, the four-derivative action (2.32) reduces to

$$S_{W}^{\text{flat}}[\phi] = -\langle \phi \,|\, \mathcal{G} \,\mathcal{I}^{-1} \,\mathcal{G} \,\phi \rangle \,, \tag{2.36}$$

which is actually the unique action invariant under both gauge symmetries as well as the Weyl transformation,

$$\delta_{\alpha} \,\phi_{\mu\nu,\rho} = \eta_{\rho[\mu} \,\alpha_{\nu]} \,. \tag{2.37}$$

This is the analogue of the Weyl gravity action, which is uniquely fixed by diffeomorphism and Weyl symmetries. Hence, one can regard  $\mathcal{S}_{W}^{\text{flat}}$  as the Weyl action for the simple hook field. In four dimensions, this action has the conformal invariance, and its form has been determined in [31] together with other two-column mixed symmetry fields.

Coming back to the AdS action  $\mathcal{S}_{W}^{\Lambda}$  of (2.32), one may wonder whether it also admits a Weyl symmetry:

$$\delta^{\Lambda}_{\alpha} \, \phi_{\mu\nu,\rho} = g^{\Lambda}_{\rho[\mu} \, \alpha_{\nu]} \,. \tag{2.38}$$

One can check the invariance of (2.32) under (2.38) by a brute force computation, but there is in fact a simpler way to see it. Since any linear combination of gauge symmetries should be a gauge symmetry, taking the following gauge parameters,

$$\epsilon_{\mu\nu}(\alpha) = \nabla_{(\mu} \alpha_{\nu)}, \qquad \theta_{\mu\nu}(\alpha) = -\frac{1}{3} \nabla_{[\mu} \alpha_{\nu]}, \qquad (2.39)$$

we immediately get the Weyl transformation,

$$\left(\delta_{\epsilon(\alpha)} + \delta_{\theta(\alpha)}\right)\phi_{\mu\nu,\rho} = \frac{\Lambda}{(d-1)(d-2)} g_{\rho[\mu}^{\Lambda} \alpha_{\nu]}. \tag{2.40}$$

This is due to the cancellation of derivative terms. Hence, in a sense, the Weyl symmetry arises as the (A)dS remnant of the gauge-for-gauge symmetry in flat space (2.12). We conclude that any theory of hook field in (A)dS, which is invariant with respect to both gauge symmetries is also Weyl invariant.

This conclusion can be generalized to any mixed symmetry fields in (A)dS. Therefore, any theory, that is invariant with respect to two distinct gauge transformations of mixed-symmetry field, has to enjoy certain Weyl symmetry.

This argument is true even for symmetric fields. Let us take the example of spintwo field  $\phi_{\mu\nu}$ : if there exists a theory of  $\phi_{\mu\nu}$  invariant with respect to both massless and partially-massless gauge symmetries,

$$\delta\phi_{\mu\nu} = \nabla_{(\mu}\epsilon_{\nu)} + \nabla_{\mu}\nabla_{\nu}\sigma + \frac{2\Lambda}{(d-1)(d-2)}g^{\Lambda}_{\mu\nu}\sigma, \qquad (2.41)$$

then such a theory will also admit a Weyl symmetry. For the demonstration, it is enough to set the gauge parameter  $\epsilon_{\mu} = -\nabla_{\mu}\sigma$  to get  $\delta\phi_{\mu\nu} = \frac{2\Lambda}{(d-1)(d-2)} g_{\mu\nu}^{\Lambda} \sigma$ . The action having both of massless and partially-massless gauge symmetries can be constructed analogously to (2.32), using the massless Einstein operator  $\mathcal{G}_{m}^{\Lambda}$  and the partially-massless one  $\mathcal{G}_{pm}^{\Lambda}$ , as

$$S_{\mathbf{W}}^{\Lambda}[\phi] = -\langle \phi \,|\, \mathcal{G}_{m}^{\Lambda} \,\mathcal{I}_{\mathrm{FP}}^{-1} \,\mathcal{G}_{nm}^{\Lambda} \,\phi \rangle \,, \tag{2.42}$$

where  $\mathcal{G}_{pm}^{\Lambda}$  differs  $\mathcal{G}_{m}^{\Lambda}$  by a particular mass term given through a Fierz-Pauli operator  $\mathcal{I}_{\text{FP}}$ . This four-derivative action (2.42) is nothing but Conformal Gravity linearized around (A)dS background. As the reader might notice, this is the inverse logic of the discussion in [18]. In general, one can use this argument to support the conjecture of [19] about the spectrum of higher-derivative Weyl-like actions for symmetric higher spin fields.

#### 2.3.1 Curvature Formulation

In order to grasp some geometrical intuitions, let us revisit our construction in terms of (generalized) curvatures. First of all, notice that an analogue of the Riemann curvature can be constructed for the hook field. It is a tensor having the symmetry of the GL(d) Young diagram given in flat space by

$$(\mathcal{R}\,\phi)_{\mu_1\mu_2\mu_3,}{}^{\nu_1\nu_2} = 6\,\partial^{[\nu_1}\,\partial_{[\mu_1}\,\phi_{\mu_2\mu_3],}{}^{\nu_2]}\,,\tag{2.43}$$

and is invariant with respect to both gauge symmetries (2.9). Around (A)dS background, it is impossible to construct a curvature invariant under both gauge transformations (2.18). Instead, we can consider the curvature,

$$(\mathcal{R}^{\Lambda}_{\square} \phi)_{\mu_1 \mu_2 \mu_3}^{\nu_1 \nu_2} = 6 \nabla^{[\nu_1} \nabla_{[\mu_1} \phi_{\mu_2 \mu_3]}^{\nu_2]}, \qquad (2.44)$$

which is invariant under the gauge transformation  $\delta_{\epsilon}^{\Lambda}$  with symmetric parameter. In order to get a curvature that is invariant with respect to the gauge symmetry  $\delta_{\theta}^{\Lambda}$  with antisymmetric parameter, one should deform the above curvature into

$$(\mathcal{R}_{\exists}^{\Lambda} \phi)_{\mu_1 \mu_2 \mu_3,}{}^{\nu_1 \nu_2} = (\mathcal{R}_{\Box}^{\Lambda} \phi)_{\mu_1 \mu_2 \mu_3,}{}^{\nu_1 \nu_2} + \frac{24 \Lambda}{(d-1)(d-2)} \delta_{[\mu_1}^{[\nu_1} \phi_{\mu_2 \mu_3],}{}^{\nu_2]}. \tag{2.45}$$

These two curvatures can be directly related to the Einstein actions (2.25): the Einstein tensors (2.20) and (2.22) are given through the curvatures (2.44) and (2.45) respectively as

$$(\mathcal{G}_{\square}^{\Lambda}\phi)_{\mu\nu,}{}^{\rho} = -\frac{1}{12} \, \delta_{\mu\nu\alpha\beta}^{\rho\gamma\delta\sigma} \, (\mathcal{R}_{\square}^{\Lambda}\phi)_{\gamma\delta\sigma,}{}^{\alpha\beta} \,, \qquad (\mathcal{G}_{\mathbb{H}}^{\Lambda}\phi)_{\mu\nu,}{}^{\rho} = -\frac{1}{12} \, \delta_{\mu\nu\alpha\beta}^{\rho\gamma\delta\sigma} \, (\mathcal{R}_{\mathbb{H}}^{\Lambda}\phi)_{\gamma\delta\sigma,}{}^{\alpha\beta} \,, \quad (2.46)$$

in the same manner as Einstein tensor is given through the Riemann curvature: remind, that Einstein tensor can be written as  $G^{\nu}_{\mu} = R^{\nu}_{\mu} - \frac{1}{2} \, \delta^{\nu}_{\mu} \, R = -\frac{1}{4} \, \delta^{\nu\alpha\beta}_{\mu\lambda\rho} \, R_{\alpha\beta,}^{\quad \lambda\rho}$ , where  $R_{\alpha\beta,\lambda\rho}$  is Riemann curvature,  $R_{\mu\nu}$  and R are Ricci curvature and its trace respectively. Hence, it is clear how each curvature is related to the two-derivative Einstein action for the simple hook field.

Let us now move on to the four-derivative action  $\mathcal{S}_W^{\Lambda}$ , which admits Weyl symmetry. In the case of gravity, the Weyl gravity action is simply given by the square of Weyl tensor. Hence, the action  $\mathcal{S}_W^{\Lambda}$  may also admit such a description. We can first define the Weyl tensor from the curvature as the traceless part of the latter. This can be conveniently done by introducing the traceless projector  $\mathcal{T}$  whose explicit form is not necessary for now. In terms of  $\mathcal{T}$ , the Weyl tensor can be determined as

$$W = \mathcal{T} \mathcal{R}_{\square}^{\Lambda} = \mathcal{T} \mathcal{R}_{H}^{\Lambda}. \tag{2.47}$$

Remark that the definition of Weyl tensor is the same for both curvatures  $R_{\square}^{\Lambda}$  and  $R_{\square}^{\Lambda}$  because the difference between two curvatures is precisely a trace term which is projected out under the action of  $\mathcal{T}$ . Therefore, the Weyl tensor is invariant under both gauge symmetries. As the invariance with respect to both gauge symmetries implies the invariance under Weyl symmetry transformation, the tensor  $(\mathcal{W}\,\phi)_{\mu\nu\rho,\lambda\kappa}$  can be rightfully referred to as Weyl tensor. Again, mimicking the Weyl gravity case, we can consider a four-derivative action from the square of the Weyl tensor. Since the four-derivative action invariant under both gauge symmetries is unique, they should be simply related with a proportionality constant as

$$S_{\mathbf{W}}^{\Lambda}[\phi] = -\frac{d-3}{d-4} \langle \mathcal{W} \phi | \mathcal{W} \phi \rangle.$$
 (2.48)

The uniqueness of the four derivative action with both symmetries in (A)dS can be checked directly, but can be also understood by the following simple argument: As we have shown above, any action for hook field with at most four derivatives and invariant with respect to both gauge symmetries in (A)dS enjoys Weyl symmetry, and therefore coincides with the action (2.36) in the flat limit. It follows then, that if there are distinct actions with those symmetries, their difference is encoded in two-derivative terms proportional to (A)dS curvature. Therefore there is a linear combination of them that is a two-derivative action invariant under both symmetries. Since there is no such two-derivative action, it follows that the four-derivative action (2.32) is unique.

In order to show that the formula (2.48) coincides with (2.32), one needs to perform several integrations by part ignoring the boundary terms. In the case of Weyl gravity, this amounts to

$$W^{\mu\nu,\rho\lambda} W_{\mu\nu,\rho\lambda} = \mathcal{L}_{GB} + \frac{4(d-3)}{d-2} \left( R^{\mu\nu} R_{\mu\nu} - \frac{d}{4(d-1)} R^2 \right), \qquad (2.49)$$

where the Gauss-Bonet (GB) term given by

$$\mathcal{L}_{GB} = \frac{1}{4} \, \delta^{\mu_1 \, \mu_2 \, \mu_3 \, \mu_4}_{\nu_1 \, \nu_2 \, \nu_3 \, \nu_4} \, R_{\mu_1 \, \mu_2},^{\nu_1 \nu_2}_{\nu_2} \, R^{\nu_3 \nu_4},_{\mu_3 \mu_4} \,. \tag{2.50}$$

vanishes identically in  $d \leq 3$  and is a topological invariant in d = 4. In higher dimensions, its linearization always reduces to a boundary term. Here again, the 'Weyl squared' action for the simple hook field can be related to the form (2.32) by adding a GB-like term for the simple hook field. In flat space, such term is given by

$$\mathcal{L}_{GB}^{flat}(\phi) = \frac{1}{2! \, 3!} \, \delta_{\nu_1 \, \nu_2 \, \nu_3 \, \nu_4 \, \nu_5}^{\mu_1 \, \mu_2 \, \mu_3 \, \mu_4 \, \mu_5} \, (\mathcal{R} \, \phi)_{\mu_1 \, \mu_2 \, \mu_3}^{\nu_1 \, \nu_2} \, (\mathcal{R} \, \phi)^{\nu_3 \, \nu_4 \, \nu_5}_{\mu_4 \, \mu_5} \,. \tag{2.51}$$

This term now vanishes identically in  $d \le 4$  and becomes a boundary term in  $d \ge 5$ . There is a unique (A)dS generalization of (2.51) that is invariant with respect to both gauge symmetries and reduces to a boundary term in  $d \ge 5$ . It is given by

$$\mathcal{L}_{GB}^{\Lambda}(\phi) = \frac{1}{2! \, 3!} \, \delta_{\nu_1 \, \nu_2 \, \nu_3 \, \nu_4 \, \nu_5}^{\mu_1 \, \mu_2 \, \mu_3 \, \mu_4 \, \mu_5} \, (\mathcal{R}_{\square}^{\Lambda} \, \phi)_{\mu_1 \mu_2 \mu_3}^{\nu_1 \nu_2} \, (\mathcal{R}_{\square}^{\Lambda} \, \phi)^{\nu_3 \nu_4 \nu_5}_{\mu_4 \mu_5} \,. \tag{2.52}$$

The Weyl squared action (2.48) differs from the action (2.32) by the boundary term (2.52):

$$\langle \mathcal{W} \, \phi \, | \, \mathcal{W} \, \phi \rangle = \frac{1}{12} \int d^d x \sqrt{|g^{\Lambda}|} \, \mathcal{L}_{GB}^{\Lambda}(\phi) - \frac{d-4}{d-3} \, \langle \phi \, | \, \mathcal{G}_{\Box}^{\Lambda} \mathcal{I}^{-1} \, \mathcal{G}_{\dot{\Box}}^{\Lambda} \, \phi \rangle \,. \tag{2.53}$$

Let us notice that both the Weyl tensor and the GB-like term (2.51) vanish identically for  $d \leq 4$ , while the action (2.32) does not vanish in d=4. The equation (2.53) still holds, since the coefficient in front of the last term vanishes for d=4. In fact, in four dimensions, the action (2.32) describes special spectra, which is different from the off-shell field. We will come back to this later.

#### 2.3.2 Factorization and Degrees of Freedom

The Lagrangian (2.32) can be written in the ordinary-derivative form by introducing an auxiliary field  $f_{\mu\nu,\rho}$  as

$$S_{\mathbf{W}}^{\Lambda}[\phi] \simeq m_{\Lambda}^{2} \left( \langle \phi \mid \mathcal{G}_{\square}^{\Lambda} \phi \rangle + 2 \langle \phi \mid \mathcal{G}_{\square}^{\Lambda} f \rangle + m_{\Lambda}^{2} \langle f \mid \mathcal{I} f \rangle \right). \tag{2.54}$$

We can solve the equation for f and go back to the Lagrangian (2.32). The Lagrangian (2.54) can be diagonalized into

$$S_{\mathrm{W}}^{\Lambda}[\phi] \simeq m_{\Lambda}^{2} \left( S_{\mathrm{Em}}^{\Lambda}[h] - S_{\mathrm{EH}}^{\Lambda}[f] \right).$$
 (2.55)

using the field redefinition,

$$\phi_{\mu\nu,\rho} = h_{\mu\nu,\rho} - f_{\mu\nu,\rho} \,. \tag{2.56}$$

This rewriting procedure is singular in the flat limit (note that the action (2.54) is not well defined in the flat space limit  $m_{\Lambda}^2 \to 0$ ), while for the action (2.32) the flat space limit is well defined and the number of DoFs of the theory is the same in (A)dS and flat spaces. The action (2.55) contains both short representations of the hook in (A)dS<sub>d</sub>. In the flat-space limit, it propagates two relatively ghost hook-helicity  $\square$  modes (which are propagating only for  $d \geq 5$ ), supplemented with a spin-two helicity mode  $\square$  and a two-form helicity mode  $\square$ . Even in (A)dS background, the reasoning in terms of the SO(d-2) helicity modes can serve as a useful guideline, despite the fact that they are not irreducible anymore when  $\Lambda \neq 0$ .

Let us now analyze the DoFs of the theory  $\mathcal{S}_W^{\Lambda}$  in terms of the helicity modes. The result of the analysis can be schematically formulated as

$$m_{\Lambda}^{2}\left(\begin{array}{c} \\ \\ \\ \\ \end{array}\right) + \Lambda \begin{array}{c} \\ \\ \\ \\ \end{array}\right) - m_{\Lambda}^{2}\left(\begin{array}{c} \\ \\ \\ \\ \end{array}\right) - \Lambda \begin{array}{c} \\ \\ \\ \end{array}\right).$$
 (2.57)

The h-field comes with a supplementary propagating two-form mode  $\Box_h$ , which is unitary in dS space and non-unitary in AdS [27]. Since the action for that mode comes with the factor  $m_{\Lambda}^2$  in front, which is positive in AdS and negative in dS, corresponding two-form  $\Box_h$  is non-unitary in both cases, while the leading hook  $\Box_f$  itself is unitary in AdS and non-unitary in dS. In the same way, the spin-two mode  $\Box_f$  of the f-field has the same sign of kinetic term as the leading hook  $\Box_f$  in AdS and opposite sign in dS, therefore is non-unitary in both cases, taking into account the factor  $-m_{\Lambda}^2$  in front of the action for f. We conclude, that among the four propagating helicity modes, only one of the hook modes has positive sign of kinetic term, therefore the theory is non-unitary for any  $d \geq 5$ , when hook modes propagate.

In four dimensions, the helicity group is SO(2), which does not have any hook representation, therefore the hook helicity modes do not propagate, hence we are left with a pseudo scalar  $\Box_h \sim \bullet_h$  and a spin-two  $\Box_f$  modes both non-unitary. We can render these two modes into unitary one by simply introducing a negative sign factor in the original action. In this way, we can describe a unitary system of spin-two and scalar in four dimensions using the four-derivative Weyl action  $\mathcal{S}_W^{\Lambda}$ . This theory is unitary in four dimensions and has the same spectrum for any value of  $\Lambda$  (including zero), and coincides with the massless limit of 4d New Massive Gravity [23] in flat space. This action can be equally obtained by a particular dimensional reduction of massless spin-two field from 5d down to 4d [24].

#### 2.3.3 Massive Deformation

Let us consider mass deformations of the Weyl action with Einstein terms. Again we have two options: the first one is to introduce  $S_{Em}^{\Lambda}$  term so that the total action preserves the magain gauge symmetry:

$$\mathcal{S}_{\square}^{\text{massive}}[\phi, m^2] = \mathcal{S}_{W}^{\Lambda}[\phi] + m^2 \mathcal{S}_{E\square}^{\Lambda}[\phi], \qquad (2.58)$$

and the other possibility is keeping the B symmetry by considering.

$$\mathcal{S}_{\mathbb{H}}^{\text{massive}}[\phi, m^2] = \mathcal{S}_{\mathbf{W}}^{\Lambda}[\phi] + m^2 \mathcal{S}_{\mathbf{E}\mathbb{H}}^{\Lambda}[\phi]. \tag{2.59}$$

In each cases, we can introduce an auxiliary field f and do a proper redefinition of the type ' $\phi = h - f$ ' as in the Weyl action case so that the four-derivative actions turn into a two-derivative ones,

$$\mathcal{S}_{\square}^{\text{massive}}[\phi, m^2] \simeq \left( m_{\Lambda}^2 + m^2 \right) \left( \left\langle h \mid \mathcal{G}_{\square}^{\Lambda} h \right\rangle - \left\langle f \mid \left( \mathcal{G}_{\mathbb{H}}^{\Lambda} - m^2 \mathcal{I} \right) f \right\rangle \right), \tag{2.60}$$

$$\mathcal{S}_{\mathbb{H}}^{\text{massive}}[\phi, m^2] \simeq (m_{\Lambda}^2 - m^2) \left( \langle h \mid (\mathcal{G}_{\square}^{\Lambda} - m^2 \mathcal{I}) h \rangle - \langle f \mid \mathcal{G}_{\mathbb{H}}^{\Lambda} f \rangle \right). \tag{2.61}$$

One can see that depending on which symmetry we decide to preserve, the other field — whose gauge symmetry is spoiled by the Einstein term — acquires a mass. Let us also note that the action  $\mathcal{S}_{\square/\mathbb{H}}^{\text{massive}}$  becomes singular when  $\mp m^2$  approaches  $m_{\Lambda}^2$ . This corresponds to the coincidence limit of two spectra described by h and f.

Now let us focus on the four dimensional case. In the previous section, we have already remarked that the Weyl action  $\mathcal{S}_W^{\Lambda}$  describes a massless spin-two and a pseudo scalar in four dimensions. With the deformation of the Einstein term, the action  $\mathcal{S}_{\square}^{\text{massive}}$  describes

$$S_{\square}^{\text{massive}} \sim \left( m_{\Lambda}^2 + m^2 \right) \left( \Lambda \bullet_h - \boxed{\qquad}_f^{SO(3)} \right),$$
 (2.62)

whereas the  $\mathcal{S}_{\mathbb{H}}^{\mathrm{massive}}$  gives

$$\mathcal{S}_{\mathbb{H}}^{\text{massive}} \sim \left( m_{\Lambda}^2 - m^2 \right) \left( \square_h^{SO(3)} + \Lambda \square_f^{SO(2)} \right),$$
 (2.63)

where we have dualized all the modes except for the last one  $\Box_f^{SO(2)}$  and used the fact that the hook helicity mode identically vanishes in four dimensions. Let us remark that in  $\mathcal{S}_{\Box}^{\text{massive}}$  the two modes — scalar  $\bullet_h$  and massive spin two  $\Box_f^{SO(3)}$  — have the same sign of the kinetic term only for  $\Lambda < 0$ , namely in AdS space. On the contrary, the two modes of  $\mathcal{S}_{\Box}^{\text{massive}}$  — massive and massless spin two fields  $\Box_h^{SO(3)}$  and  $\Box_f^{SO(2)}$  — have the same sign in dS space  $(\Lambda > 0)$ . Hence, again by introducing an overall minus sign in these actions, we obtain two models of four-derivative theories with unitary propagation. It will be interesting to explore possible links of these four-derivative formulations with existing models. For instance, the field content of  $\mathcal{S}_{\Box}^{\text{massive}}$  coincides with that of bigravity [32] proposed by Hasan and Rosen.

# 3 Generalizations

All the discussions about the simple hook field can be straightforwardly generalized to more general fields having the symmetry of two-column Young diagrams. The number of columns is playing the role of 'spin' and the two-column fields show many common features with spin-two field. In particular, their Weyl actions contain four derivatives. This considerably reduces the technical complexities and allows us to perform the analysis explicitly.

The only qualitative difference of two-columns fields with respect to the hook one rises in the case where the two columns have the same height, which we refer to as 'long window' diagram. Hence, we shall first consider the non-window case in below and then do a separate analysis for the long window case in the succeeding section.

#### 3.1 Two Columns

In this section, we will generalize the results of Section 2 for a tensor field  $\phi_{\mu_1\cdots\mu_p}$ , having the symmetry of two column Young diagram:

$$\phi_{\mu_1\cdots\mu_p}, \begin{array}{c} \nu_1\cdots\nu_q \\ \\ \end{array} \sim \left[ \begin{array}{c} q \\ \\ \end{array} \right], \tag{3.1}$$

where we assume p is strictly larger than q.

#### **Einstein Action** 3.1.1

The Labastida action for this field is given making use of the diagram,

which is made of two fields and two derivatives contracted with the generalized Kronecker delta  $\delta^{\nu_1\cdots\nu_{p+q+1}}_{\mu_1\cdots\mu_{p+q+1}}$ . The corresponding Einstein tensor is hence given by

$$(\mathcal{G}\,\phi)_{\mu_1\cdots\mu_p,}{}^{\nu_1\cdots\nu_q} = -\frac{(-1)^{(p+1)(q+1)}}{p!\,q!}\,\delta^{\nu_1\cdots\nu_{p+q+1}}_{\mu_1\cdots\mu_{p+q+1}}\,\partial^{\mu_{p+1}}\,\partial_{\nu_{q+1}}\,\phi_{\nu_{q+2}\cdots\nu_{p+q+1}},{}^{\mu_{p+2}\cdots\mu_{p+q+1}}\,, \tag{3.3}$$

which is invariant with respect to two gauge symmetries:

$$\delta_{\epsilon} \phi_{\mu_1 \cdots \mu_p},^{\nu_1 \cdots \nu_q} = \partial_{[\mu_1} \epsilon_{\mu_2 \cdots \mu_p]},^{\nu_1 \cdots \nu_q}, \tag{3.4}$$

$$\delta_{\epsilon} \, \phi_{\mu_{1} \cdots \mu_{p}},^{\nu_{1} \cdots \nu_{q}} = \partial_{[\mu_{1}} \, \epsilon_{\mu_{2} \cdots \mu_{p}]},^{\nu_{1} \cdots \nu_{q}},$$

$$\delta_{\theta} \, \phi_{\mu_{1} \cdots \mu_{p}},^{\nu_{1} \cdots \nu_{q}} = \partial^{[\nu_{1}} \, \theta_{\mu_{1} \cdots \mu_{p}},^{\nu_{2} \cdots \nu_{q}]} + (-1)^{p+1} \, \frac{p}{p-q+1} \, \partial_{[\mu_{1}} \, \theta_{\mu_{2} \cdots \mu_{p}]},^{[\nu_{1}, \nu_{2} \cdots \nu_{q}]},$$

$$(3.4)$$

with gauge parameters of the following Young symmetry type:

$$\epsilon_{\mu_1\cdots\mu_{p-1}}, \stackrel{\nu_1\cdots\nu_q}{\sim} \sim \stackrel{p-1}{ } \boxed{\qquad} , \qquad \theta_{\mu_1\cdots\mu_p}, \stackrel{\nu_1\cdots\nu_{q-1}}{\sim} \sim \boxed{\qquad} p \boxed{\qquad} .$$
 (3.6)

The Lagrangian (3.2), or equivalently the Einstein tensor (3.3) can be also given through the gauge invariant curvature tensor  $(\mathcal{R} \phi)_{\mu_1 \cdots \mu_{p+1}}$ ,  $\nu_1 \cdots \nu_{q+1}$ , which again has a symmetry of two-column Young diagram but with additional boxes in each column (corresponding to derivatives):

$$(\mathcal{R}\,\phi)_{\mu_1\cdots\mu_{p+1}}, \stackrel{\nu_1\cdots\nu_{q+1}}{\sim} \sim \begin{bmatrix} p & & & & \\ p & & & \\ & \overline{\partial} & & & \\ & & \overline{\partial} & & & \\ & & & \overline{\partial} & & \\ & & \overline{\partial} & & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \\ & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} & \overline{\partial} &$$

Its explicit expression reads

$$(\mathcal{R}\,\phi)_{\mu_1\cdots\mu_{p+1}},^{\nu_1\cdots\nu_{q+1}} = (p+1)(q+1)\,\partial_{[\mu_1}\,\partial^{[\nu_1}\,\phi_{\mu_2\cdots\mu_{p+1}]},^{\nu_2\cdots\nu_{q+1}]}\,,\tag{3.8}$$

and it is invariant with respect to both gauge symmetries (3.5).

Around (A)dS background, the curvature cannot be gauge invariant with respect to both of the symmetries, and one has to choose a zero-derivative deformation to preserve one of the symmetries:

$$\delta_{\epsilon}^{\Lambda} \phi_{\mu_1 \cdots \mu_p},^{\nu_1 \cdots \nu_q} = \nabla_{[\mu_1} \epsilon_{\mu_2 \cdots \mu_p]},^{\nu_1 \cdots \nu_q}, \tag{3.9}$$

$$\delta_{\epsilon}^{\Lambda} \phi_{\mu_{1} \cdots \mu_{p}, \nu_{1} \cdots \nu_{q}} = \nabla_{[\mu_{1}} \epsilon_{\mu_{2} \cdots \mu_{p}], \nu_{1} \cdots \nu_{q}}, \qquad (3.9)$$

$$\delta_{\theta}^{\Lambda} \phi_{\mu_{1} \cdots \mu_{p}, \nu_{1} \cdots \nu_{q}} = \nabla^{[\nu_{1}} \theta_{\mu_{1} \cdots \mu_{p}, \nu_{2} \cdots \nu_{q}]} + (-1)^{p+1} \frac{p}{p-q+1} \nabla_{[\mu_{1}} \theta_{\mu_{2} \cdots \mu_{p}]}^{[\nu_{1}, \nu_{2} \cdots \nu_{q}]}, \quad (3.10)$$

The curvature invariant under the  $\delta_{\epsilon}^{\Lambda}$  symmetry takes the form,

$$(\mathcal{R}^{\Lambda}_{[p-1,q]}\phi)_{\mu_1\cdots\mu_{p+1},}{}^{\nu_1\cdots\nu_{q+1}} = (p+1)(q+1)\nabla^{[\nu_1}\nabla_{[\mu_1}\phi_{\mu_2\cdots\mu_{p+1}],}{}^{\nu_2\cdots\nu_{q+1}]}, \tag{3.11}$$

which is again a simple extension of the flat space curvature (3.8). Considering now the curvature invariant under the  $\delta_{\theta}^{\Lambda}$  transformation, we find

$$\begin{aligned}
&(\mathcal{R}^{\Lambda}_{[p,q-1]}\phi)_{\mu_{1}\cdots\mu_{p+1},}^{\nu_{1}\cdots\nu_{q+1}} \\
&= (\mathcal{R}^{\Lambda}_{[p-1,q]}\phi)_{\mu_{1}\cdots\mu_{p+1},}^{\nu_{1}\cdots\nu_{q+1}} - \frac{2(p-q+1)\Lambda}{(d-1)(d-2)} \delta^{[\nu_{1}}_{[\mu_{1}}\phi_{\mu_{2}\cdots\mu_{p+1}],}^{\nu_{2}\cdots\nu_{q+1}]} \\
&= (p+1)(q+1)\nabla_{[\mu_{1}}\nabla^{[\nu_{1}}\phi_{\mu_{2}\cdots\mu_{p+1}],}^{\nu_{2}\cdots\nu_{q+1}]} - \frac{2\Lambda}{(d-1)(d-2)} \delta^{[\nu_{1}}_{[\mu_{1}}\phi_{\mu_{2}\cdots\mu_{p+1}],}^{\nu_{2}\cdots\nu_{q+1}]} . (3.12)
\end{aligned}$$

We can see that these two curvatures differ by a zero-derivative term. One can then construct Einstein tensors out of these curvatures as

$$(\mathcal{G}^{\Lambda}_{[p-1,q]}\phi)_{\mu_1\cdots\mu_p},^{\nu_1\cdots\nu_q} = -\frac{(-1)^{(p+1)(q+1)}}{(p+1)!(q+1)!}\delta^{\nu_1\cdots\nu_{p+q+1}}_{\mu_1\cdots\mu_{p+q+1}}(\mathcal{R}^{\Lambda}_{[p-1,q]}\phi)_{\nu_{q+1}\cdots\nu_{p+q+1}},^{\mu_{p+1}\cdots\mu_{p+q+1}},(3.13)$$

$$(\mathcal{G}^{\Lambda}_{[p,q-1]}\phi)_{\mu_{1}\cdots\mu_{p}},^{\nu_{1}\cdots\nu_{q}} = -\frac{(-1)^{(p+1)(q+1)}}{(p+1)!(q+1)!}\delta^{\nu_{1}\cdots\nu_{p+q+1}}_{\mu_{1}\cdots\mu_{p+q+1}}(\mathcal{R}^{\Lambda}_{[p,q-1]}\phi)_{\nu_{q+1}\cdots\nu_{p+q+1}},^{\mu_{p+1}\cdots\mu_{p+q+1}}.(3.14)$$

The Einstein tensors immediately define the Einstein actions  $\mathcal{S}^{\Lambda}_{\mathrm{E}\,[p-1,q]}$  and  $\mathcal{S}^{\Lambda}_{\mathrm{E}\,[p,q-1]}$  for the two-column field (3.1). The action  $\mathcal{S}_{\mathrm{E}\,[p-1,q]}^{\Lambda}$  is (not) unitary around (A)dS and the opposite for  $\mathcal{S}^{\Lambda}_{E[p,q-1]}$ . An example of unitary Einstein action is that of the [d-2,1] field in d dimensions, for which the  $\mathcal{S}^{\Lambda}_{E[d-2,0]}$  action has a propagating degree of freedom of a dual graviton. This action was studied recently in [17]. This same field is used in New Massive Gravity in arbitrary dimensions [24, 25]. The relation between these models will become transparent in the next section.

#### 3.1.2 Weyl Action

Let us now move to the Weyl action. For that, we consider the Weyl tensor for two column field (3.1) by making use of the trace projector  $\mathcal{T}$  as

$$W = \mathcal{T} \mathcal{R}^{\Lambda}_{[p-1,q]} = \mathcal{T} \mathcal{R}^{\Lambda}_{[p,q-1]}. \tag{3.15}$$

The form of the trace projector acting on a [p+1, q+1] tensor can be conveniently given by the following expression:

$$\mathcal{T} = \frac{(-1)^{(p+1)(q+1)}}{(p+1)!(q+1)!} \left( \delta_{p+q+2} - \frac{d-p-q-1}{d-p-q} \delta_{p+q+1} (\delta_{p+q})^{-1} \delta_{p+q+1} \right), \tag{3.16}$$

where the operator  $\delta_n$  acts on the field  $f_{\mu_1\cdots\mu_p}$ ,  $\nu_1\cdots\nu_q$  of type [p,q] in the following way

$$(\delta_n f)_{\nu_{q+1}\cdots\nu_n},^{\mu_{p+1}\cdots\mu_n} = \delta^{\mu_1\cdots\mu_n}_{\nu_1\cdots\nu_n} f_{\mu_1\cdots\mu_p},^{\nu_1\cdots\nu_q},$$
(3.17)

contracting all of its indices to the generalized Kronecker-delta. Inverse operator  $(\delta_{p+q})^{-1}$  is defined as:

$$(\delta_{p+q}(\delta_{p+q})^{-1}f)_{\mu_1\dots\mu_p},^{\nu_1\dots\nu_q} = f_{\mu_1\dots\mu_p},^{\nu_1\dots\nu_q} = ((\delta_{p+q})^{-1}\delta_{p+q}f)_{\mu_1\dots\mu_p},^{\nu_1\dots\nu_q}.$$
(3.18)

The tracelessness of the expression (3.16) can be proven using the identity,

$$\delta_{\nu_1}^{\mu_1} \, \delta_{\mu_1 \dots \mu_n}^{\nu_1 \dots \nu_n} = (d - n + 1) \, \delta_{\mu_2 \dots \mu_n}^{\nu_2 \dots \nu_n} \,. \tag{3.19}$$

It is again clear that Weyl tensor (3.15) is invariant under both gauge symmetries (3.10) hence also under Weyl transformation:

$$\delta_{\alpha} \,\phi_{\mu_1 \cdots \mu_p,}^{\nu_1 \cdots \nu_q} = \delta^{[\nu_1}_{[\mu_1} \,\alpha_{\mu_2 \cdots \mu_p]}^{\nu_2 \cdots \nu_q]} \,. \tag{3.20}$$

Analogously to the hook case, the Weyl action is given by the square of Weyl tensors,

$$\mathcal{S}_{\mathbf{w}}^{\Lambda}[\phi] = -\frac{(-1)^{p+q} (d-p-q)}{d-p-q-1} \langle \mathcal{W} \phi | \mathcal{W} \phi \rangle, \qquad (3.21)$$

where the coefficient is fixed such that the Weyl action takes another representation,

$$S_{\mathbf{W}}^{\Lambda}[\phi] = -\langle \phi \,|\, \mathcal{G}_{[p-1,q]}^{\Lambda} \,\mathcal{I}^{-1} \,\mathcal{G}_{[p,q-1]}^{\Lambda} \,\phi \rangle \,, \tag{3.22}$$

up to a GB-like total derivative term,

$$\mathcal{L}_{GB}^{\Lambda}(\phi) = (-1)^{(p+1)(q+1)} \left\langle \mathcal{R}_{[p-1,q]}^{\Lambda} \phi \, | \, \delta_{p+q+2} \, \mathcal{R}_{[p,q-1]}^{\Lambda} \, \phi \right\rangle. \tag{3.23}$$

In (3.22) the mass operator  $\mathcal{I}$  is defined as

$$\mathcal{I} = \frac{(-1)^{p\,q}}{p!\,q!}\,\delta_{p+q}\,. \tag{3.24}$$

Hence, one can see that the entire construction of the hook example can be generalized to the two column case in a straightforward manner. The DoF can be conventiently analyzed by using the same factorization technique as the simple hook case. Skipping the straightforward derivation part, let us directly present the end result,

which is the natural generalization of the result (2.57). All these DoF vanish in dimensions  $d \leq p+q$ . When d=p+q+1, the helicity mode [p,q] vanishes, but there remains propagating DoFs given by

where we have dualized both modes. Note that these modes have the same kinetic term sign, hence can describe unitary propagation by introducing a negative factor in the original action. Notice that when q = 1 we get in this way the DoF of a scalar and a helicity two mode. The corresponding [p, q] = [d - 2, 1] Weyl action coincides in fact with the massless limit of New Massive Gravity action in any d [24].

#### 3.2 Long Window

In the special case where the height of two columns are equal, that is q = p:

$$\phi_{\mu_1\cdots\mu_p}, \begin{array}{c} \nu_1\cdots\nu_p \\ \\ \end{array} \sim \left[ \begin{array}{c} p \\ \\ \end{array} \right], \qquad (3.27)$$

the analysis is no more analogous to the hook field case, but actually more similar to the spin-two case. Let us see explicitly how this works. First of all, the field  $\phi_{\mu_1\cdots\mu_p}$ ,  $\nu_1\cdots\nu_p$  admits only the gauge symmetry generated by the parameter of [p,p-1] Young diagram, because there is no Young diagram [p-1,p]. Therefore, even in flat space, the long window has only one gauge symmetry.

In (A)dS background, we can first consider the curvature  $\mathcal{R}^{\Lambda}_{[p,p-1]}$  (3.12) or equivalently the Einstein tensor  $\mathcal{G}^{\Lambda}_{[p,p-1]}$  having the [p,p-1] gauge symmetry. In this case, there is nothing different from the generic two column case and we can obtain the corresponding tensor and action.

Considering now the curvature  $\mathcal{R}^{\Lambda}_{[p-1,p]}$  (3.11) or the Einstein tensor  $\mathcal{G}^{\Lambda}_{[p-1,p]}$ , we first note that they cannot have the [p-1,p] gauge symmetry as it simply does not exist. Instead, one may wonder whether this action still plays a special role. It tuns out that

with the sacrifice of the [p-1,p] gauge symmetry, the action  $\mathcal{S}^{\Lambda}_{\mathrm{E}[p-1,p]}$  acquires a new gauge symmetry,

$$\delta_{\epsilon}^{\Lambda} \phi_{\mu_1 \cdots \mu_p},^{\nu_1 \cdots \nu_p} = \left( \nabla_{[\mu_1} \nabla^{[\nu_1} - \frac{2\Lambda}{(d-1)(d-2)} \delta^{[\nu_1]}_{[\mu_1} \right) \epsilon_{\mu_1 \cdots \mu_{p-1}},^{\nu_1 \cdots \nu_{p-1}}, \tag{3.28}$$

with the gauge parameter,

$$\epsilon_{\mu_1 \cdots \mu_{p-1}}, \qquad p-1 \qquad \qquad p-1 \qquad \qquad p-1$$
 (3.29)

This is clearly reminiscent of the partially-massless gauge symmetry. Indeed, when p = 1 the corresponding action coincides with that of partially-massless spin two. Now considering the [p-1, p-1] gauge symmetry (3.28), we can construct a one-derivative gauge invariant, or curvature, as

$$(\mathcal{C}\,\phi)_{\mu_1\cdots\mu_{p+1}},^{\nu_1\cdots\nu_p} = (p+1)\,\nabla_{[\mu_1}\,\phi_{\mu_2\cdots\mu_{p+1}]},^{\nu_1\cdots\nu_p},\tag{3.30}$$

and the action  $\mathcal{S}^{\Lambda}_{\mathrm{E}[p-1,p]}$  can be also expressed as the square of this curvature and its traces. We can proceed to construct a four-derivative action  $\mathcal{S}^{\Lambda}_{\mathrm{W}}$  having both of [p,p-1] and [p-1,p-1] gauge symmetries:

$$\delta^{\Lambda} \phi_{\mu_{1} \cdots \mu_{p}, \nu_{1} \cdots \nu_{p}} = \left( \nabla_{[\mu_{1}} \nabla^{[\nu_{1}} - \frac{2\Lambda}{(d-1)(d-2)} \delta^{[\nu_{1}}_{[\mu_{1}}) \epsilon_{\mu_{1} \cdots \mu_{p-1}, \nu_{1} \cdots \nu_{p-1}} + \right. \\ + \left. \nabla^{[\nu_{1}} \theta_{\mu_{1} \cdots \mu_{p}, \nu_{2} \cdots \nu_{p}]} + (-1)^{p+1} p \nabla_{[\mu_{1}} \theta_{\mu_{2} \cdots \mu_{p}]}^{[\nu_{1}, \nu_{2} \cdots \nu_{p}]}. \right. (3.31)$$

Such an action will be automatically invariant under the Weyl transformation (3.20) because it can be realized as a particular configuration of (3.31). The action  $\mathcal{S}_{\mathbf{W}}^{\Lambda}$  can be constructed exactly in the same way as in the generic [p,q] case, either using Weyl tensor as in (3.21) or using Einstein tensors as in (3.22).

The DoF of the Weyl action can be analysed in terms of SO(d-2) representations in a similar manner as before. The result reads

$$m_{\Lambda}^{2} \left( \begin{array}{c|c} & & & \\ & + \Lambda & \\ & & \\ h & & \end{array} \right) - m_{\Lambda}^{2} \left( \begin{array}{c} & \\ & \\ f & \end{array} \right), \tag{3.32}$$

Notice that the h- and f-fields are the long-window analogs of partially-massless spin two and massless spin two. When 2p = d - 1, we end up with only one mode,

$$\begin{array}{c|c}
p-1 & & \\
& & \\
h & & \\
\end{array} \tag{3.33}$$

which we have dualized. Notice that when p=1, this action describes a scalar mode in three-dimension. This is nothing but the propagating content of the 3d parity-invariant linear Weyl gravity — or massless limit of New Massive Gravity [22] — and the scalar mode corresponds to the parity-invariant partially-massless spin two. It would be also interesting to remark that in 5d, we get in this way a helicity two mode from the action of the [2, 2] window field, which might provide an alternative theory of Gravity.

# 3.3 Massive Deformation and New Massive Gravity

The massive deformation of the Weyl action for the generic two-column fields follows the same pattern of the simple hook case:

$$\mathcal{S}_{[p,q-1]}^{\text{massive}}[\phi, m^{2}] = \mathcal{S}_{W}^{\Lambda}[\phi] + m^{2} \mathcal{S}_{E[p,q-1]}^{\Lambda}[\phi] 
\simeq (m_{\Lambda}^{2} + m^{2}) \left( \langle h | \mathcal{G}_{[p-1,q]}^{\Lambda} h \rangle - \langle f | \left( \mathcal{G}_{[p,q-1]}^{\Lambda} - m^{2} \mathcal{I} \right) f \rangle \right), \qquad (3.34) 
\mathcal{S}_{[p-1,q]}^{\text{massive}}[\phi, m^{2}] = \mathcal{S}_{W}^{\Lambda}[\phi] + m^{2} \mathcal{S}_{E[p-1,q]}^{\Lambda}[\phi] 
\simeq (m_{\Lambda}^{2} - m^{2}) \left( \langle h | \left( \mathcal{G}_{[p-1,q]}^{\Lambda} - m^{2} \mathcal{I} \right) h \rangle - \langle f | \mathcal{G}_{[p,q-1]}^{\Lambda} f \rangle \right), \qquad (3.35)$$

which describe one massive [p,q] mode and one massless  $[p,q] \oplus [p,q-1]$  or  $[p,q] \oplus [p-1,q]$  mode, where the latter massless mode becomes partially-massless for q=p case.

Let us consider the dimensions d=p+q+1, where the leading [p,q]-helicity modes disappear leaving only lower helicity modes. In this case, there is a preferred choice of massive deformation between  $\mathcal{S}_{[p,q-1]}^{\text{massive}}$  and  $\mathcal{S}_{[p-1,q]}^{\text{massive}}$  depending on the sign of the cosmological constant as in the simple hook case. In dS background, only the massive action  $\mathcal{S}_{[p,q-1]}^{\text{massive}}$  can describe massive and massless modes with the same sign of kinetic terms because

$$S_{[p,q-1]}^{\text{massive}} \sim (m_{\Lambda}^2 - m^2) \left( \begin{array}{c} q & q \\ q & \\ & \\ & h \end{array} \right) + \Lambda \left( \begin{array}{c} q & q \\ & \\ & \\ & f \end{array} \right). \tag{3.36}$$

Instead in AdS background, the other action  $\mathcal{S}_{[p-1,q]}^{\text{massive}}$  is preferred since

$$S_{[p-1,q]}^{\text{massive}} \sim (m_{\Lambda}^2 + m^2) \left( \Lambda \right) \left( \Lambda \right) \left( \frac{q-1}{1 - q} \right) \left( \frac{$$

Notice that in the above formulas two different Young diagrams — one carrying SO(d-2) representation and the other carrying SO(d-1) representation — are used.

Now taking the long-window case (2p = d - 1), we have two analogue massive theories,

Hence, the massive action  $\mathcal{S}_{[p-1,p-1]}^{\text{massive}}$  having the 'partially-massless' gauge symmetry have one additional [p-1,p-1] helicity mode compared to the action  $\mathcal{S}_{[p,p-1]}^{\text{massive}}$  corresponding to 'massless' gauge symmetry. The action  $\mathcal{S}_{[p,p-1]}^{\text{massive}}$  has only single massive irreducible [p,p] mode, hence unitary in both of  $AdS_{2p+1}$  and  $dS_{2p+1}$  background. On the contrary, the other action  $\mathcal{S}_{[p-1,p]}^{\text{massive}}$  propagates two modes with the same kinetic term sign only in AdS background. Focusing on the p=1 case, the action  $\mathcal{S}_{[1,0]}^{\text{massive}}$  describes a massive spin two in three dimensions and actually coincides with the linearization of the New Massive Gravity [22]. We can actually see here, that in  $AdS_3$  there is an alternative "New Massive Gravity", that makes use of the partially-massless symmetry instead of the diffeomorphism one, and contains additional scalar in the spectrum.

The term "New Massive Gravity" in higher dimensions refers to the models with fields of type [p,1] (see [24, 25] for related discussion) in dimensions d=p+2. Let us note, that the flat limit of  $(A)dS_{p+2}$  New Massive Gravity in dimensions higher than three is not smooth. In flat limit we have only a massive spin two mode, while in AdS we have unitary model with (massive spin two + massless scalar), while in dS the unitary model contains (massive spin two + massless spin two). Instead, in three dimensions, one can have New Massive Gravity with the same spectrum around flat and constantly curved backgrounds, at least at the linearized level.

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