# Collision induced broadening of $\nu_1$ band and ground state spectral lines of sulfur dioxide perturbed by $N_2$ and $O_2$

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#### **Abstract**

To monitor the constituents and trace pollutants of Earth atmosphere and understand its evolution, accurate spectroscopic parameters are fundamental information.  $SO_2$  is produced by both natural and anthropogenic sources and it is one of the principal causes of acid rains as well as an important component of fine aerosol particles, once oxidized to sulfate. The present work aims at determining  $SO_2$  broadening parameters using  $N_2$  and  $O_2$  as atmospherically relevant damping gases. Measurements are carried out in the infrared (IR) and mm-/sub-mm wave regions, around 8.8  $\mu$ m and in the 104 GHz – 1.1 THz interval, respectively. IR ro-vibrational transitions are recorded by using a tunable diode laser spectrometer, whereas the microwave spectra are recorded by using a frequency-modulated millimeter-/submillimeter-wave spectrometer.  $SO_2$ - $N_2$  and  $SO_2$ - $O_2$  collisional cross sections are retrieved for several  $\nu_1$  band ro-vibrational transitions of  ${}^{32}SO_2$ , for some transitions belonging to either  $\nu_1 + \nu_2 - \nu_2$  of  ${}^{32}SO_2$  or  $\nu_1$  of  ${}^{34}SO_2$  as well as for about 20 pure rotational transitions in the vibrational ground state of the main isotopic species. From  $N_2$ - and  $O_2$ -broadening coefficients the broadening parameters of  $SO_2$  in air are derived. The work is completed with the study of the dependence of foreign broadening coefficients on the rotational quantum numbers.

**Keywords:** SO<sub>2</sub>; atmospheric monitoring; N<sub>2</sub>-, O<sub>2</sub>-, air-broadening parameters; infrared spectroscopy; microwave spectroscopy.

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#### Introduction

Nowadays, remote sensing techniques allow monitoring and accurately retrieving the concentration profiles of atmospheric constituents and trace pollutants, since these molecules have strong rotation and vibration-rotation absorptions in the microwave (MW) and infrared (IR) spectral domains, respectively [1-4]. The satellites used for sounding the terrestrial atmosphere embark spectrometers that provide a large amount of spectral information at ever increasing quality in terms of spectral coverage, resolution and signal-to-noise ratio [1]. For these reasons, spectroscopic parameters are of fundamental importance for exploiting remote sensing applications in atmospheric and climate research, environmental monitoring and gas-phase analysis. As a matter of fact, only their accurate knowledge allows an accurate retrieval of concentrations and distributions of the gas phase molecular species in the atmosphere. The relevant spectroscopic parameters include line-by-line parameters, i.e., transition frequencies and their intensities, pressure broadening coefficients and pressure induced shifts, and their temperature dependence. The existing spectroscopic data are then collected into a number of different databases, among which the most important ones are HITRAN [5], GEISA [6], JPL [7], and the Cologne database [8]. Within this framework, the aim of laboratory spectroscopy is to provide spectroscopic parameters for a wide variety of species of atmospheric, astrophysical and industrial relevance. Furthermore the study of collisional broadening and shifting coefficients, these being related to the intermolecular potential, can shed light on the driving forces ruling the scattering events in the gas phase (see e.g. Refs. 9-13 and references therein).

Sulfur dioxide (SO<sub>2</sub>) is an important molecule for Earth atmosphere since it actively enters in the sulfur cycle and it is one of the causes of acid rains. SO<sub>2</sub> has a lifetime in the atmosphere of about a day, it is oxidized quickly in the atmosphere, thus leading to aerosol formation and acid deposition [14]. As sulfate, it is an important component of fine aerosol particles (PM10 and PM2.5). Furthermore, sulfate aerosol also affect Earth's radiation balance either thorough direct scattering of sunlight or indirectly via modification of cloud albedoes. On a global scale, the majority of natural SO<sub>2</sub> is produced by volcanoes [15-17] and by the oxidation of sulfur gases produced by the decomposition of plants. Therefore, natural emissions usually occur at high altitude or far from city centers, and hence the SO<sub>2</sub> background concentration in clear air is about 1 ppb [18]. On the other hand, a sizeable amount of SO<sub>2</sub> is of anthropogenic origin, being produced by combustion of fossil fuels as well as by non-ferrous smelting processes for the conversion of ores to free metals [18]. In addition, this compound is largely employed in food-preserving and wine

making industries. It is thus not surprising that SO<sub>2</sub> has received a great deal of attention from the scientific community and it is still the subject of a number of spectroscopic studies.

The fundamental  $v_1$  and  $v_3$  bands of  ${}^{32}S^{16}O_2$  isotopologue have been thoroughly analyzed in 1980s by Guelachvili and co-workers [19,20] and subsequently reinvestigated by Flaud et al. together with the  $v_2$  fundamental and the  $2v_2$  -  $v_2$  hot band [21]. In 2005 Müller and Brünker have accurately re-determined the rotational parameters for the ground and  $v_2 = 1$  states [22]. Some rovibrational transitions, belonging to  $v_1$  and  $v_3$  bands, have been object of study by Sumpf with the aim of determining line-by-line transition intensities [23], while accurate intensity determinations have been performed within the spectral region 940 - 1400 cm<sup>-1</sup> region by Fourier transform IR spectroscopy [24]. Since 1992 different studies have focused on the determination of SO<sub>2</sub> self-[23,25-32] and foreign-broadening coefficients [26,27,33-37]. A few years ago, a complete listing of sulfur dioxide self-broadening coefficients has been compiled by combining the results obtained from IR and MW spectroscopy with semiclassical calculations [38]. In addition, the homodimer of SO<sub>2</sub> has been investigated by Tasinato and Grimme, theoretically using dispersion-corrected density functional theory (DFT-D3) as well as experimentally by means of tunable diode laser (TDL) IR spectroscopy [39]. In that work, the dissociation energies of (SO<sub>2</sub>)<sub>2</sub> and (CH<sub>2</sub>F<sub>2</sub>)<sub>2</sub> have been determined experimentally from the broadening of the ro-vibrational transitions of the corresponding monomers collisionally perturbed by a range of damping gases [39]. Recently, SO<sub>2</sub> has been included in an HITRAN-like database of line parameters for molecules of planetary interest perturbed by H<sub>2</sub>, He or CO<sub>2</sub> [40]. Very recently, Ceselin at al. exploited IR and mm/sub-mm wave spectroscopy to retrieve new line-by-line pressure broadening parameters of SO<sub>2</sub> perturbed by He, H<sub>2</sub> and CO<sub>2</sub> [41].

In the present contribution, our work aiming at the determination of  $SO_2$  broadening parameters for atmospheric and astrochemical applications is extended by considering the atmospherically relevant  $N_2$  and  $O_2$  buffer gases.

# **Experimental details**

The broadening parameters of SO<sub>2</sub> perturbed by N<sub>2</sub> and O<sub>2</sub> as buffer gases have been determined both in the infrared (IR) and millimeter/sub-millimeter (mm-/sub-mm) wave region. IR measurements have been carried out at the Laboratory of Molecular Spectroscopy of Venice (LMS-Ve), whereas mm-/sub-mm ones have been performed in the Laboratory of mm/sub-mm wave Spectroscopy of Bologna (LMS-Bo).

For what concerns IR measurements, SO<sub>2</sub> high resolution spectra were recorded by using a TDL spectrometer installed at LMS-Ve and working in a three beam configuration [42], in which the radiation emitted from a lead salt laser was split into three beams. The main part of the radiation was directed through the SO<sub>2</sub> sample contained in either a 92.3 cm or 68.1 cm path length cell. The other two beams were sent to a Germanium etalon and to the SO<sub>2</sub> reference cell, respectively, both employed for calibration purposes. Each beam was collected by its own HgCdTe detector, cryogenically cooled at liquid nitrogen temperature. The acquired signals were digitized by a four channel ultrafast acquisition card with a 14 bit vertical resolution [43].

The spectra were recorded at  $(296 \pm 1)$  K by using two different lasers centered at 8.1 and 8.9 µm, respectively, both cooled at liquid nitrogen temperature. Up to 1024 independent scans were averaged to generate each spectrum. Pressure broadening measurements were carried out by perturbing a fixed amount of SO<sub>2</sub> around 100 Pa (according to the intensities of the spectral lines) and adding increasing quantities of damping gas up to a total pressure of 45.40 hPa at most. The pressure was measured by using two different vacuum gauges with a full scale range of 10 and 100 hPa, respectively, each with a quoted manufacture accuracy of 0.15% on the full scale. Two independent series of measurements were performed in order to increase the accuracy of the retrieved pressure broadening parameters [43,44]. An elapsed time of 10 – 15 min was adopted between the filling of the sample cell and the recording of the spectra in order to promote gas homogenization inside the cell. The SO<sub>2</sub> gas was supplied by Sigma - Aldrich with a purity of 99.9%, whereas N<sub>2</sub> and O<sub>2</sub> damping gases were provided by SIAD with a purity greater than 99.9% and 99.5% respectively. Line frequencies that derive from high resolution Fourier Transform IR data ware used for wavenumber calibration of the acquired spectra [37]. The accuracy of the calibration was estimated around  $3.3 \times 10^{-4}$  cm<sup>-1</sup>. Absorbance spectra were obtained from the incident and transmitted radiation intensities according to the Beer-Lambert's law, with the incident intensity retrieved by fitting no absorption regions (i.e. the baseline) to a polynomial function.

The lineshape analysis was performed by using the Visual LineShape Fitting program (VLSFP) running in multiline mode [45]. Line-by-line parameters were obtained by fitting all the absorption lines in a given micro-window to the Voigt profile (VP) and the spectra recorded at different pressures were fitted independently from each other. Concerning the two lasers adopted in the present work, the one centered at 8.1  $\mu$ m presented an instrumental lineshape function that could be well reproduced as a Lorentzian function, whereas that emitting around 8.9  $\mu$ m was characterized by a Gaussian instrumental function. In the first case during the fits employing the VP model, Gaussian half widths were fixed at the molecular contribution  $\gamma_{\rm D}^{\rm so_2}$  due to the thermal motion

$$\gamma_{\rm D}^{\rm SO_2} = \frac{v_{if}}{c} \left( \frac{2k_{\rm B} N_a T \ln 2}{m} \right)^{1/2} \tag{1}$$

where  $v_{ij}$  is the frequency of the  $f \leftarrow i$  ro-vibrational transition, c represents the speed of light,  $k_B$  is the Boltzman's constant,  $N_a$  is the Avogadro's constant, T is the absolute temperature and m is the mass of the radiating species. The Lorentzian instrumental half width was considered as described in the next section (see equation 3). In the case of a Gaussian instrumental function, the Doppler half width was fixed to an effective value given by the following equation

$$\gamma_{\mathrm{D}}^{\mathrm{eff}} = \left[ \left( \gamma_{\mathrm{D}}^{\mathrm{SO}_2} \right)^2 + \left( \gamma_{\mathrm{D}}^{\mathrm{TDL}} \right)^2 \right]^{1/2} \tag{2}$$

where  $\gamma_D^{TDL}$  is the instrumental Doppler half width. The recording of the low-pressure (< 30 Pa) spectra of pure SO<sub>2</sub> and the subsequent fitting of the spectral features to a Gaussian profile provided the effective Doppler half widths with typical values around  $1.15 \times 10^{-3}$  cm<sup>-1</sup>. The assignments of the recorded spectral features were performed by using the HITRAN database [5].

Concerning the MW measurements, the recording of pure rotational spectra of  $SO_2$  perturbed by either  $N_2$  or  $O_2$  in the mm-/sub-mm wave region was performed using a frequency-modulated computer-controlled spectrometer working in the 65 GHz - 1.6 THz frequency range, with the 104 GHz - 1.1 THz range actually considered. The detailed description of the spectrometer is provided in Refs. [46-48]. To briefly summarize, the radiation source is an interchangeable system consisting of Gunn-diode oscillators and Gunn-diode-driven frequency multipliers, phase-locked to a rubidium frequency standard. The frequency modulation is achieved by sine-wave modulating the 72 MHz local oscillator of the synchronization loop at 1.666 kHz and, by tuning the Lock-in amplifier to twice the modulation frequency, second harmonic detection is performed. The spectrometer is equipped with a liquid-He InSb detector for THz measurements and Schottky detectors for recordings below 500 GHz.

In this work, a previous investigation of the pressure broadening of rotational lines of SO<sub>2</sub> perturbed by N<sub>2</sub> and O<sub>2</sub> was extended. To investigate the dependence of pressure-broadening parameters on the J and  $K_a$  quantum numbers, the following  $\Delta J = 0$  and  $\Delta K_a = +1$  transitions were considered [36]. The first set consists of transitions with the same starting value for  $K_a$  ( $K_a = 11 \leftarrow 10$ ) in order to investigate the J dependence:  $J = 11_{11,1} \leftarrow 11_{10,2}$ ,  $J = 12_{11,1} \leftarrow 12_{10,2}$ ,  $J = 15_{11,5}$ 

 $\leftarrow 15_{10,6}, J = 16_{11,5} \leftarrow 16_{10,6}, J = 26_{11,15} \leftarrow 26_{10,16}, J = 28_{11,17} \leftarrow 28_{10,18}, J = 34_{11,23} \leftarrow 34_{10,24}, J = 42_{11,31} \leftarrow 42_{10,32}, J = 51_{11,41} \leftarrow 51_{10,42}, J = 60_{11,49} \leftarrow 60_{10,50}$ . In the second set, transitions with the same J value (J = 10) were considered in order to study the  $K_a$  dependence:  $J = 10_{1,9} \leftarrow 10_{0,10}, J = 10_{2,8} \leftarrow 10_{1,9}, J = 10_{4,6} \leftarrow 10_{3,7}, J = 10_{5,5} \leftarrow 10_{4,6}, J = 10_{6,4} \leftarrow 10_{5,5}, J = 10_{7,3} \leftarrow 10_{6,4}$ . In the present study, further measurements for these two sets of transitions were performed in order to confirm and improve the results of Ref. [36]. To extend the study, one  $\Delta J = -1$  transition,  $J = 20_{7,13} \leftarrow 21_{6,16}$ , and two  $\Delta J = +1$  transitions,  $J = 43_{5,39} \leftarrow 42_{6,36}$  and  $J = 3_{3,1} \leftarrow 2_{2,0}$ , were also considered.

The rotational lines were recorded using starting pressures of  $SO_2$  in the range 0.04 - 0.53 Pa (according to transition intensity) and then adding increasing quantities of either  $N_2$  or  $O_2$ , from ~6.7 Pa up to ~133.3 Pa. The pressure inside the cell was measured by means of a Baratron gauge with a measurable pressure range of 0.01 - 133.3 Pa. The gas pressure was continuously monitored during the measurements and it was found stable within less than 0.01 Pa, the latter value thus providing the accuracy of its measurements. All spectra were recorded at room temperature (298  $\pm$  1 K). For each transition, series of measurements were performed in order to have a total of 25 - 30 points.

To retrieve the collisional (also denoted Lorentzian) line width, the second harmonic profiles were fitted to a lineshape model that explicitly accounts for frequency modulation [49-51]. To correctly reproduce the lineshapes recorded at low pressure (up to a total pressure in the range 26.7 - 33.3 Pa), and in particular, to account for narrowing effects, the Speed Dependent Voigt Profile (SDVP) [52,53] was employed. The suitability of this line profile model is demonstrated by the small residuals (observed-calculated differences) observed, as shown in Figure 1, which provides an example of line profile analysis for some measurements of the  $J=43_{5,39} \leftarrow 42_{6,36}$  rotational transition broadened by oxygen.

#### **Results and discussion**

The radiating species investigated in the present work is an asymmetric near-prolate rotor belonging to the  $C_{2\nu}$  symmetry point group. It has three vibrational normal modes:  $\nu_1$  and  $\nu_3$  correspond to the O=S=O symmetric and asymmetric stretching, respectively, while  $\nu_2$  represents the bending motion. The  $\nu_1$  fundamental belongs to the  $A_1$  symmetry species and it gives rise to a B-type band located at about 1151.7 cm<sup>-1</sup>. The presence of two identical oxygen nuclei (zero nuclear spin) allows, for all totally symmetric states, only rotational levels with an even value of  $K_a^{"} + K_c^{"}$ .

Concerning the work carried out for retrieving foreign-broadening coefficients of  $SO_2$  perturbed by  $N_2$  and  $O_2$ , as mentioned above, the line profile analysis in the IR region was performed by using the Voigt profile model. Figure 2 presents the  $SO_2$  absorptions in the spectral micro-window between 1226.07 and 1226.78 cm<sup>-1</sup> perturbed by  $O_2$  together with the corresponding residuals referring to total pressures between 145.1 and 3179 Pa. The good residuals shown in this figure provide the evidence that the Voigt profile well reproduces the observed features within the experimental set-up and conditions employed for the present measurements. In order to retrieve  $SO_2$  foreign-broadening coefficients  $\gamma_X^0(X=N_2,O_2)$ , the collisional half widths at the temperature T,  $\Gamma_L(T)$ , as obtained from the lineshape analysis of the ro-vibrational transitions, were fitted against the partial pressure of  $N_2$  or  $O_2$  to a straight line represented by the following equation

$$\Gamma_L(T) = \Gamma_0(T) + \gamma_X^0(T) \times P_X \tag{3}$$

where  $\Gamma_0(T)$  takes into account the self-broadening contribution to the line width and the Lorentzian component of the instrumental function (if needed) and  $P_X$  represents the partial pressure of damping gas. Since slightly different quantities of  $SO_2$  were introduced into the sample cell for the two different series of measurements, to combine the results of the linear fits obtained from the two repetitions, we employed the following procedure [43,44,54,55]. For each spectral line we subtracted the value of the intercept  $\Gamma_0$  from each  $\Gamma_L$ , thus obtaining the so-called normalized Lorentzian half width  $\Gamma_{\rm NL}$ . A second fit involving all normalized half widths was then performed to retrieve the final value of the pressure broadening coefficient according to the following equation

$$\Gamma_{NL} = \gamma_X^0 \times P_X \tag{4}$$

An example of the linear fits involving normalized Lorentzian half widths obtained for the rovibrational transition  $29_{16,14} \leftarrow 28_{15,13}$  centered at 1221.29402 cm<sup>-1</sup> and perturbed by N<sub>2</sub> and O<sub>2</sub> is shown in Figure 3. Table 1 collects the obtained SO<sub>2</sub>-N<sub>2</sub> and SO<sub>2</sub>-O<sub>2</sub> broadening coefficients for the  $\nu_1$  band transitions of the  $^{32}S^{16}O_2$  isotopologue, whereas Table 2 lists the results for the  $\nu_1 + \nu_2 - \nu_2$ hot band of  $^{32}S^{16}O_2$  as well as those for the  $\nu_1$  fundamental band of  $^{34}S^{16}O_2$ . In fact, in addition to the transitions belonging to the  $\nu_1$  band of the main isotopic species, it was possible to assign and analyze a number of lines due to either the hot band or the  $\nu_1$  fundamental of  $^{34}S^{16}O_2$ . In these tables, the SO<sub>2</sub>-air broadening parameters ( $\gamma_{air}^0$ ) are also reported, and they were derived from the measured N<sub>2</sub>- and O<sub>2</sub>- broadening coefficients by means of the following equation

$$\gamma_{\rm air}^0 = 0.79 \gamma_{\rm N_2}^0 + 0.21 \gamma_{\rm O_2}^0. \tag{5}$$

As it can be seen, for the  $v_1$  band of  $^{32}S^{16}O_2$ , the recorded transitions span the quantum number range  $14 \le J'' \le 56$ ,  $3 \le K_a'' \le 21$  and they belong to the P, Q or R branches. It should be noted that, for three transitions of the  $v_1$  band of  $^{32}S^{16}O_2$  ( $24_{8,16} \leftarrow 25_{9,17}$ ;  $19_{13,7} \leftarrow 19_{14,6}$ ;  $18_{13,5} \leftarrow 18_{14,4}$ ) and one of  $^{34}S^{16}O_2$  ( $25_{11,15} \leftarrow 25_{12,14}$ ) it has been possible to determine only  $N_2$ -broadening coefficients as from  $N_2$  to  $O_2$  broadening experiments, the laser emission profile shifted slightly so that it was not possible to exactly recover the same spectral region.

The dependence of the N<sub>2</sub>- and O<sub>2</sub>- broadening coefficients of  $^{32}S^{16}O_2$  on the  $K_a^{"}$  (pseudo-) quantum number for  $v_1$  band transitions within each branch is illustrated in Figure 4. In this figure, the outcomes from our previous investigation [37] are also reported, showing that the present determinations confirm those findings. Concerning the P-branch transitions, the N<sub>2</sub>-broadening parameters do not show any clear trend being almost uniformly scattered around 0.0958 cm<sup>-1</sup>atm<sup>-1</sup>. On the other hand, SO<sub>2</sub>-O<sub>2</sub> collisional cross sections appear to slightly increase up to  $K_a^{"}=11$  and then they remain almost constant. Moving to the Q-branch, both N2- and O2-broadening coefficients keep, over the  $K_a^{"}$  range here considered (13  $\leq K_a^{"} \leq 21$ ), an almost constant value around 0.0946 and 0.0717 cm<sup>-1</sup>atm<sup>-1</sup>, respectively. As far as the R-branch transitions are concerned, the very small dependence on  $K_a^{"}$  is confirmed, even though that  $N_2$  collisional cross sections seem to first decrease up to  $K_a$  = 14, 15 and then they tend to increase slightly as the quantum number raises further. The trend observed for O2-broadening parameters suggests a very small increase with increasing  $K_a^{"}$  values. The fact that in the R-branch a small dependence of the broadening coefficients on  $K_a$  seems to emerge is probably due to the smaller uncertainties affecting these transitions. Overall, the obtained N2- and O2- broadening coefficients for the three branches present an average value of  $(0.0945 \pm 0.0059)$  cm<sup>-1</sup>atm<sup>-1</sup> and  $(0.0716 \pm 0.0047)$  cm<sup>-1</sup>atm<sup>-1</sup>, respectively.

At this point a comment concerning the comparison with the available literature data is deserved. As already pointed out, the broadening parameters here determined are in close agreement with those found by Tasinato et al. [37], and they also well match with the N<sub>2</sub>- and O<sub>2</sub>-

broadening coefficients retrieved by Cazzoli et al. [36], Sumpf and collaborators [26,27] and Ball et al. [33].

Concerning the analysis in the MW region, for a given transition, for each series of measurements, the pressure-broadening coefficient  $\gamma_X^0$  at 298 K,  $\gamma_X^0$  (298), was derived by a linear fit of the collisional half widths,  $\Gamma_L(T)$ , against the perturber partial pressure  $P_X$ , as detailed in Eq. (3), with the half widths weighted according to the reciprocal of the squared uncertainties obtained from the profile analysis. To put together the data coming from the different series of measurements carried out for each transition, a procedure analogous to that described above for IR measurements and introduced in Ref. [56] was employed to obtain the normalized Lorentzian half widths,  $\Gamma_{NL}$ (298). Subsequently, a second fit involving these normalized half widths was performed in order to obtain the final value of the pressure-broadening parameter  $\gamma_X^0$  (298) for the transition under consideration, see Eq. (4). Figure 5, which depicts the linear fit of the normalized collisional half widths for the  $J = 20_{7,13} \leftarrow 21_{6,16}$  (258.7 GHz) transition, provides a typical example, thus showing how well the collisional half widths follow a linear behavior against the pressure.

The  $N_2$ - and  $O_2$ -broadening coefficients of the rotational transitions considered are collected in Table 3 together with the corresponding quantum numbers and transition frequencies. In Table 3, the air-broadening coefficients, obtained by means of Eq. (5) are also given. Strictly speaking, the first 16 lines listed in this Table were already investigated in Ref. [36] and, as already pointed out, they have been further investigated in this work.

From the inspection of the values of Table 3, it is noted that the pressure-broadening coefficients vary slightly upon variation of either J'' or  $K_a^*$ , the largest variations being of about 10 - 15%. For example, within the set of transitions selected for investigating the J'' dependence, the maximum variation observed is ~10%, despite the large range of J'', from 11 to 60, considered. The very limited dependence of the pressure broadening coefficients on the J'' or  $K_a^*$  quantum numbers was already pointed out [36], even if polynomial expressions were there deduced to model the variations observed. For the sake of completeness, the  $K_a^*$  and J'' dependence of the broadening parameters is graphically illustrated in Figure S.1 of the Supplementary Information (SI). A clear trend cannot be derived in agreement with the findings reported for the  $\nu_1$  band. To the best of our knowledge, there is not available in the literature any other study reporting either J'' or  $K_a^*$  dependence of rotational transitions. However, it deserves to be mentioned the investigation of the temperature dependence carried out by Ball et al. [33], who investigated the 90 – 500 K range. If one compares the pressure-broadening coefficients determined at room temperature for the four

rotational transitions there studied, it is noted that the values are in line with those derived in the present work:  $J = 18_{3,15} - 18_{2,16}$  (240.247 GHz):  $\gamma_{N_2}^0 = 0.1072(54)$  cm<sup>-1</sup> atm<sup>-1</sup>,  $\gamma_{O_2}^0 = 0.0735(37)$  cm<sup>-1</sup> atm<sup>-1</sup>,  $J = 13_{1,13} - 12_{0,12}$  (251.199 GHz):  $\gamma_{N_2}^0 = 0.0974(49)$  cm<sup>-1</sup> atm<sup>-1</sup>,  $\gamma_{O_2}^0 = 0.0814(41)$  cm<sup>-1</sup> atm<sup>-1</sup>,  $J = 4_{2,2} - 3_{1,3}$  (235.152 GHz):  $\gamma_{N_2}^0 = 0.1202(60)$  cm<sup>-1</sup> atm<sup>-1</sup>,  $\gamma_{O_2}^0 = 0.0920(46)$  cm<sup>-1</sup> atm<sup>-1</sup>,  $J = 26_{4,22} - 26_{3,23}$  (280.807 GHz):  $\gamma_{N_2}^0 = 0.1067(53)$  cm<sup>-1</sup> atm<sup>-1</sup>,  $\gamma_{O_2}^0 = 0.0776(39)$  cm<sup>-1</sup> atm<sup>-1</sup>.

For the sake of completeness, even if it is not of relevance for the compilation of databases finalized at providing data for remote sensing of the Earth atmosphere, the speed dependence of the relaxation rate can be derived from line profile analyses performed using the SDVP model. As expected, the latter contribution shows a linear dependence on pressure and can be fitted with a linear expression analogous to Equations (3) and (4) to derive the so-called  $\gamma_2$  parameter [12,53,57]. However, even if the expected linear behavior was confirmed in all cases, very small values, about  $10^{-6}$  -  $10^{-5}$  cm<sup>-1</sup>atm<sup>-1</sup>, were obtained (see also Ref. [36]). This outcome further supports the decision to not report any results, neither graphically nor in table.

The overall trend, against  $K_a^n$ , obtained for N<sub>2</sub>- and O<sub>2</sub>-broadening coefficients of both  $\nu_1$  band- and ground state-transitions is presented in Figure 6. It is interesting to note that, actually, the range of quantum numbers accessed in the IR and MW regions is complementary: ground state transitions mostly involve  $K_a^n$  values from 0 to 6 and several transitions characterized by  $K_a^n = 10$  and different values of  $J^n$  in order to access their  $J^n$  dependence. Conversely, IR measurements span the  $K_a^n$  range from 7 to 21. Some variations with the quantum number seem to take place for  $K_a^n \leq 8$ , and then no evident trend can be reported. This tends to suggest that more detailed information may be gained if it was possible to study the behavior of collisional cross sections within each  $J^n$  and  $J^n$  and  $J^n$  sub-branch, as we performed for SO<sub>2</sub> self-broadening parameters [38]. Under this point of view, the support coming from theoretical calculations would be very helpful in order to get deeper insights.

At this point one may wonder whether some differences between the broadening parameters obtained at LMS-Ve and LMS-Bo can be due to the different line-shape models employed or to a vibrational dependence of the collisional cross sections. The latter effect can be excluded, as previously pointed out in the study of self-broadening coefficients [38]. A possible vibrational dependence may arise when going to higher quanta of vibrational excitation. Concerning the different line-shape models employed, the broadening parameters obtained by using the SDVP which accounts for collisional narrowing effects are expected to differ by only 2 - 3% from those retrieved within the VP model, as shown in Ref. [41]. Furthermore, the use of different line-shape

functions in the IR and MW regions is well justified by the different experimental conditions employed. In fact, for IR measurements, the pressures employed are, in general, two orders of magnitude greater than those used for MW experiments. At such higher pressures narrowing effects are washed out by the collisional broadening. Rather, the differences between MW and IR collisional parameters may be ascribed to the different experimental set-up adopted in the corresponding laboratories. On the other hand, it should also be pointed out that the small differences between MW and IR broadening parameters may be only apparent, because of the limited J values covered in the MW region for  $0 \le K_a^* \le 5$ . Indeed, when more MW data are available, as in the case of  $K_a^* = 10$ , no evident difference is observed between the IR and MW pressure broadening coefficients.

Finally, Figure 7 depicts the SO<sub>2</sub>-air pressure broadening coefficients as a function of the  $K_a^-$  quantum number, together with a comparison to the corresponding values in the recently updated HITRAN database [58]. The overall agreement appears quite good, yet the HITRAN values are slightly larger than those here determined: while the air broadening coefficients here derived average to  $(0.0914 \pm 0.0064)$  cm<sup>-1</sup>atm<sup>-1</sup>, the corresponding ones listed in the HITRAN database have an average value of  $(0.0986 \pm 0.0036)$  cm<sup>-1</sup>atm<sup>-1</sup>.

#### **Conclusions**

In the present work, a line-by-line list of  $SO_2$  foreign-broadening coefficients has been retrieved by the analysis of several ro-vibrational transitions belonging to the  $^{32}S^{16}O_2 \nu_1$  band (in the 8.8 µm region) as well as a number of pure rotational transitions of the vibrational ground state (in the mm-/sub-mm wave range), with  $O_2$  and  $N_2$  as damping gases. In addition, 10 transitions of the  $\nu_1 + \nu_2 - \nu_2$   $^{32}S^{16}O_2$  hot band and 6 lines belonging to  $\nu_1$  of  $^{34}S^{16}O_2$  have also been analyzed in the IR region. IR experiments have been carried out within the atmospheric window by using the TDL spectrometer installed in the Laboratory of high resolution IR spectroscopy at University of Venice and the line-shape analysis has been performed by using the Voigt profile. For what concern pure rotational transitions, the spectra have been recorded in the 104 GHz – 1.1 THz frequency range by using the frequency-modulated spectrometer installed at the laboratory of mm-/sub-mm wave spectroscopy of University of Bologna. For the line-shape analysis, the speed-dependent Voigt profile has been employed. In addition, for both IR and MW transitions, from the measured  $N_2$ - and  $N_2$ -broadening coefficients, the air-broadening coefficients have been derived and compared to

those available in the HITRAN database. A fairly good agreement has been found, only a small overestimation with respect to the database entries has been noted.

Concerning the quantum number dependence of the collisional cross sections, the results obtained from the line-shape analysis of the spectra recorded in both the IR and MW regions point out a very small dependence on both  $K_a^{"}$  and  $J^{"}$ . In this context, theoretical calculations of collisional cross sections are strongly encouraged. We can conclude that the outcomes here presented suggest the need of further detailed studies about the pressure broadening parameters of SO<sub>2</sub> collisionally perturbed by different buffer gases. Under this point of view, this work represents a further extension of our project aimed at determining the broadening coefficients of SO<sub>2</sub> for atmospheric and astrochemical applications.

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### **Supplementary Information**

Figure S.1 reporting the dependence of ground state  $N_2$ - and  $O_2$ - broadening coefficients of  $SO_2$  on the rotational quantum numbers.

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# **Figure Captions**

**Figure 1.** Oxygen broadening of the 249.0 GHz line ( $J = 43_{5,39} \leftarrow 42_{6,36}$ ) of sulfur dioxide: an example of measurements (in black) analyzed by using the SDVP model (in red). The residuals (three times the observed – calculated difference values) are also depicted (in green).

**Figure 2.** SO<sub>2</sub> ro-vibrational lines (gray circles) in the spectral region between 1226.07 and 1226.78 cm<sup>-1</sup> (experimental conditions  $P_{SO_2} = 145.1$  Pa, path length = 68.1 cm, T = 296 K) fitted to the VP model (red line). Black traces, with the same scale as the spectrum, show the residuals of the fits and they refer to the following O<sub>2</sub> partial pressures (from top to bottom): 0, 430.6, 1057, 1714, 2445, 3179 Pa.

**Figure 3.** Foreign broadening of the  $29_{16,14} \leftarrow 28_{15,13}$  ro-vibrational transition at 1221.29402 cm<sup>-1</sup> of  $^{32}S^{16}O_2$   $\nu_1$  band. Upper panel. Experimental collisional half widths:  $N_2$  (blue squares) and  $O_2$  (red circles). Dashed lines are linear fits of normalized Lorentzian half widths against the  $N_2$  and  $O_2$  partial pressure. Error bars are the uncertainties retrieved from the lineshape analysis. Lower panel. Residuals between experimental and fitted values.

**Figure 4.** Trend of N<sub>2</sub>- (black squares) and O<sub>2</sub>- (red circles) broadening coefficients of  $^{32}S^{16}O_2 \nu_1$  band transitions against the  $K_a^*$  pseudo-quantum number in the (a) P-, (b) Q- and (c) R- subbranches (filled squares and circles refer to the present work; empty squares and circles are data from Ref. 37).

**Figure 5.** Nitrogen broadening of the 258.7 GHz line ( $J = 20_{7,13} \leftarrow 21_{6,16}$ ) of sulfur dioxide in the vibrational ground state: linear fit of the "normalized" Lorentzian half widths against the partial pressure of N<sub>2</sub>. Error bars are three times the uncertainties retrieved from the line profile analysis.

**Figure 6.** Dependence of  $v_1$  band and ground state  $N_2$ - (black) and  $O_2$ - (red) broadening coefficients of  ${}^{32}S^{16}O_2$  on the  $K_a^{"}$  quantum number.

**Figure 7.**  $SO_2$  - air broadening parameters as a function of  $K_a$ : present work (blue crosses and triangles); Ref. 40 (blue squares); HITRAN database, Ref. 61 (green circles). Ground state

broadening coefficients (blue triangles) determined at (298  $\pm$  1 K);  $\nu_1$  broadening coefficients (blue crosses) determined at (296  $\pm$  1 K).

Table 1. Assignment, transition frequency,  $N_2$ -,  $O_2$ -, and air- broadening coefficients of  $^{32}S^{16}O_2$   $\nu_1$  band transitions.  $^a$ 

J	$K_a$	$K_c$	J"	$K_a^{"'}$	$K_c^{"'}$	$\nu_0{}^b$	$\gamma_{\mathrm{N}_2}^{0}{}^{c}$	$\gamma_{\mathrm{O}_2}^{0}{}^{c}$	$\gamma_{ m air}^0$
34	20	14	34	21	13	1083.2225 <sub>0</sub> (1 <sub>4</sub> )	0.0815(61)	$0.075_0(5_7)$	$0.080_1(6_0)$
46	10	36	47	11	37	$1083.22930_0(2_0)^d$	$0.097_5(1_1)$	$0.065_0(5_5)$	0.0907(20)
25	13	13	26	14	12	$1083.3932_3(2_0)$	$0.091_6(1_0)$	$0.084_8(1_9)$	$0.090_2(1_2)$
55	8	48	56	9	47	$1083.4374_0(2_6)$	$0.088_6(7_4)$	$0.076_0(1_9)$	$0.086_0(6_2)$
37	12	26	38	13	25	$1083.4797_9(2_2)$	$0.087_6(1_9)$	$0.068_5(1_9)$	$0.083_6(1_9)$
27	13	15	28	14	14	$1083.4937_5(1_2)$	$0.094_8(6_8)$	$0.080_7(9_4)$	$0.091_8(7_3)$
18	15	3	19	16	4	1083.48432(24)	0.10127(89)	$0.093_9(1_9)$	$0.0997(1_1)$
24	15	9	25	16	10	$1083.51701_0(3_0)$	$0.089_8(3_3)$	$0.068_5(1_5)$	$0.085_3(2_9)$
16	15	1	17	16	2	$1083.6045_3(2_4)$	$0.0987(3_0)$	$0.072_1(9_0)$	$0.093_1(4_3)$
29	12	18	30	13	17	1083.61264(27)	$0.118_2(7_6)$	$0.080_6(7_7)$	$0.110_3(7_6)$
28	14	14	29	15	15	$1083.76369(3_1)$	$0.105_7(3_3)$	$0.073_9(6_7)$	$0.099_0(4_0)$
29	20	10	29	21	9	1083.77345(27)	$0.094_6(5_4)$	$0.067_1(1_7)$	$0.088_8(4_6)$
28	14	14	28	15	13	$1102.31102_0(8_0)$	$0.087_1(7_7)$	$0.065_4(4_8)$	$0.082_5(7_1)$
34	7	27	35	8	28	1102.3423 <sub>0</sub> (1 <sub>9</sub> )	$0.087_9(2_0)$	$0.068_5(3_3)$	$0.083_8(2_3)$
20	10	10	21	11	11	1102.38693 <sub>5</sub> (1 <sub>6</sub> )	$0.096_9(2_1)$	$0.082_5(2_1)$	$0.093_9(2_1)$
27	14	14	27	15	13	1102.40127(19)	$0.101_4(3_5)$	$0.075_0(3_2)$	$0.095_9(3_4)$
25	7	19	26	8	18	$1102.4451_6(1_2)$	$0.092_4(5_8)$	$0.067_0(2_7)$	$0.087_1(5_1)$
26	14	12	26	15	11	1102.489355(45)	$0.093_4(3_2)$	$0.070_2(1_2)$	$0.088_{5}(2_{8})$
14	10	4	15	11	5	$1102.54508_0(7_0)$	$0.083_9(6_2)$	0.0737(26)	$0.081_8(5_4)$
29	8	22	30	9	21	1102.56124(16)	$0.096_5(2_0)$	$0.072_2(1_1)$	$0.091_4(1_8)$
25	14	12	25	15	11	$1102.57416_0(5_0)$	$0.092_7(2_4)$	$0.0802_6(9_8)$	$0.090_1(2_1)$
15	11	5	16	12	4	1102.61647 <sub>3</sub> (2 <sub>3</sub> )	$0.102_6(1_9)$	$0.078_3(1_1)$	$0.097_5(1_7)$
23	14	10	23	15	9	$1102.73491_0(5_0)$	$0.091_5(4_7)$	$0.075_7(1_5)$	$0.088_2(4_0)$
24	9	15	25	10	16	1102.83402(14)	$0.090_8(3_9)$	$0.068_1(1_4)$	$0.086_0(3_4)$
21	14	8	21	15	7	1102.8833 <sub>3</sub> (3 <sub>2</sub> )	$0.103_6(5_1)$	$0.070_0(1_4)$	$0.096_5(4_3)$
20	14	6	20	15	5	1102.95225(23)	$0.090_1(1_1)$	$0.068_4(2_1)$	$0.085_5(1_3)$
15	9	7	16	10	6	1102.9936 <sub>3</sub> (3 <sub>6</sub> )	$0.091_4(4_2)$	$0.0604_9(7_7)$	$0.084_9(3_5)$
19	14	6	19	15	5	$1103.0182_0(1_9)$	$0.104_8(3_2)$	$0.0708_5(3_7)$	$0.097_7(2_6)$
33	7	27	34	8	26	1103.056138(48)	$0.091_5(2_2)$	$0.067_1(2_0)$	$0.0864(2_2)$
18	14	4	18	15	3	$1103.08108_{5}(8_{5})$	$0.098_5(2_0)$	$0.070_7(3_6)$	$0.092_7(2_3)$
19	10	10	20	11	9	$1103.08938_6(6_6)$	$0.096_0(1_8)$	$0.0806_6(1_8)$	$0.092_8(1_5)$
18	9 8	9 16	19 25	10 9	10	$1103.1237_6(1_0)$	$0.0881_9(6_6)$	$0.063_2(6_1)$	$0.082_9(1_8)$
24 19		16 7	23 19	9 14	17	$1106.143_3(1_5)$	$0.095_3(4_5)$		
19	13 13		18	14 14	6	$1106.1908_0(3_0)^e$ $1106.2530_8(2_0)$	$0.103_3(3_2)$		
	13	5 5	17	14 14	4	1106.23308(20) $1106.31193_1(6_1)$	$0.100_7(2_2)$	0.064.(2.)	$0.088_5(1_0)$
17 19	9	11	20	10	4 10	1106.368124(67)	$0.0949_3(6_8)$	$0.064_3(2_3)$ $0.077_3(1_1)$	` '
48	12	36	48	13	35	1106.37991 <sub>5</sub> (4 <sub>5</sub> )	$0.109_1(2_4)$ $0.101_7(5_5)$	, ,	$0.102_4(2_1)$ $0.096_1(5_6)$
23	7	17	24	8		, ,	` '	$0.075_1(6_1)$	
23 15	13	3	2 <del>4</del> 15	8 14	16 2	$1106.3885_{6}(1_{2})$ $1106.41997_{5}(1_{5})^{f}$	$0.086_8(8_9)$ $0.096_8(1_9)$	$0.072_5(8_8)$ $0.0701_1(2_7)$	$0.083_8(8_9)$ $0.091_2(1_6)$
13	13	3 1	13	14 14	0	1106.419975(15) 1106.4691 <sub>6</sub> (1 <sub>0</sub> )	0.0908(19) 0.0906(33)	0.07011(27) $0.0754(3_0)$	0.0912(16) $0.0874(32)$
18	8	10	14 19	9	11	1106.4691 <sub>6</sub> (1 <sub>0</sub> ) 1106.49250 <sub>5</sub> (7 <sub>5</sub> )	0.0906(53) $0.088_2(5_8)$	0.0734(30) $0.071_3(7_6)$	0.0874(52) 0.0847(62)
33	6	28	34	7	27	1106.492505(75) 1106.50768 <sub>3</sub> (1 <sub>9</sub> ) <sup>g</sup>	$0.088_2(3_8)$ $0.092_7(1_1)$	0.0713(76) $0.0670_9(6_2)$	$0.087_{3}(1_{0})$
47	12	36	47	13	35	1106.507063(19)	0.072/(11) $0.1018_5(9_0)$	$0.06769(02)$ $0.068_5(5_5)$	0.0873(10) 0.0948(19)
т/	14	50	т/	13	55	1100.010105(45)	0.10103(70)	0.0003(23)	0.0778(19)

J'	$K_a$	$K_{c}^{'}$	J"	$K_a^{"'}$	$K_c^{"}$	$\nu_0{}^b$	$\gamma_{\mathrm{N_2}}^{0}{}^{c}$	$\gamma^0_{{ m O}_2}{}^c$	$\gamma_{ m air}^0$
14	10	4	15	11	5	$1106.55602_5(5_5)$	$0.101_9(1_2)$	$0.0728(1_6)$	$0.0958(1_3)$
13	9	5	14	10	4	$1106.56840_5(2_5)$	$0.088_2(6_6)$	$0.068_0(8_2)$	$0.084_0(6_9)$
41	14	28	40	13	27	$1221.0639_7(5_0)$	$0.090_4(1_2)$	$0.068_1(1_3)$	$0.0857(1_2)$
23	17	7	22	16	6	$1221.18158_5(6_5)^h$	$0.0901_0(9_9)$	0.06927(68)	$0.0857_3(9_2)$
35	15	21	34	14	20	1221.25888(12)	$0.086_2(1_3)$	$0.0655_6(5_6)$	$0.081_9(1_1)$
29	16	14	28	15	13	$1221.2938_5(1_7)^i$	$0.0921_6(7_5)$	$0.06752(3_3)$	$0.08699(6_6)$
33	16	18	32	15	17	1221.36008(48)	$0.099_1(4_6)$	0.0687(91)	0.0927(55)
26	18	8	25	17	9	$1226.0960_2(1_0)$	$0.089_3(2_8)$	$0.071_1(1_1)$	$0.085_5(2_4)$
38	16	22	37	15	23	$1226.1134_1(1_3)$	$0.090_9(1_7)$	$0.069_3(1_0)$	$0.086_4(1_6)$
32	17	15	31	16	16	$1226.18016_3(2_1)$	$0.0987_2(6_8)$	$0.0702_8(4_4)$	$0.0927_5(6_3)$
30	18	12	29	17	13	$1226.32107_0(8_0)$	$0.101_0(8_1)$	$0.081_2(1_7)$	$0.096_8(6_8)$
21	19	3	20	18	2	1226.445519(41)	$0.094_9(1_3)$	$0.0725_3(9_4)$	$0.090_2(1_2)$
39	16	24	38	15	23	$1226.63220_0(1_0)$	$0.086_5(1_1)$	$0.0635_2(9_9)$	$0.0817(1_1)$
27	18	10	26	17	9	$1226.65480_0(7_1)$	$0.0881_{7}(5_{8})$	$0.0684_1(5_7)$	$0.0840_2(5_8)$
39	13	27	38	12	26	$1216.70535_5(4_5)$	$0.0893_5(6_2)$	$0.0670_5(7_6)$	$0.0846_7(6_5)$
33	17	17	32	16	16	1226.71967 <sub>8</sub> (4 <sub>7</sub> )	$0.0935_0(7_1)$	$0.0736_0(7_3)$	$0.0893_2(7_1)$
21	16	6	20	15	5	1216.790645(55)	$0.0941_9(6_1)$	$0.0756_2(7_3)$	$0.0902_9(6_4)$
46	12	34	45	11	35	1216.80594(12)	$0.098_5(1_3)$	$0.066_6(1_7)$	$0.091_8(1_4)$
19	17	3	18	16	2	1216.82777(13)	$0.090_9(9_3)$	$0.080_2(7_8)$	$0.0887(9_0)$

<sup>&</sup>lt;sup>a</sup> Uncertainties in parentheses are one standard deviation.

<sup>&</sup>lt;sup>b</sup> Transition frequency in cm<sup>-1</sup>.

<sup>&</sup>lt;sup>c</sup> Broadening coefficient in cm<sup>-1</sup> atm<sup>-1</sup>.

<sup>&</sup>lt;sup>d</sup> Overlapped to  $54_{19,35}$  ←  $54_{20,34}$   $\nu_1$  <sup>32</sup>S<sup>16</sup>O<sub>2</sub>.

<sup>&</sup>lt;sup>e</sup> Overlapped to  $15_{8,8} \leftarrow 16_{9,7} \ v_1^{34} S^{16} O_2$ ;  $33_{5,29} \leftarrow 34_{6,28} \ v_1 + v_2 - v_2^{32} S^{16} O_2$ .

<sup>&</sup>lt;sup>f</sup> Overlapped to  $24_{6,18}$  ←  $25_{7,19} \nu_1$   $^{34}\text{S}^{16}\text{O}_2$ .

<sup>&</sup>lt;sup>g</sup> Overlapped to  $41_{11,31} \leftarrow 41_{12,30} \nu_1 + \nu_2 - \nu_2 {}^{32}S^{16}O_2$ .

<sup>&</sup>lt;sup>h</sup> Overlapped to  $39_{15,25} \leftarrow 38_{14,24} \ \nu_1 + \nu_2 - \nu_2 \ ^{32}S^{16}O_2$ ;  $35_{22,14} \leftarrow 35_{21,15} \ \nu_1 \ ^{32}S^{16}O_2$ .

<sup>&</sup>lt;sup>i</sup> Overlapped to  $34_{22,12} \leftarrow 34_{21,13} \nu_1^{32} S^{16} O_2$ .

Table 2. Assignment, transition frequency,  $N_2$ -,  $O_2$ -, and air- broadening coefficients of  $^{32}S^{16}O_2$   $\nu_1$  +  $\nu_2$ - $\nu_2$  hot band and  $^{34}S^{16}O_2$   $\nu_1$  band.  $^a$ 

J'	$K_a$	$K_{c}^{'}$	J"	$K_a^{"}$	$K_c^{"'}$	ν <sub>0</sub> <sup>b</sup>	$\gamma_{\mathrm{N}_{2}}^{0}{}^{c}$	$\gamma_{\mathrm{O}_2}^{0}{}^{c}$	$\gamma_{ m air}^{0}$
18	15	3	19	16	4	1083.48432(24)	0.10127(89)	0.0939(19)	$0.099_{7}(1_{1})$
27	13	15	28	14	14	$1083.4937_5(1_2)$	$0.094_8(6_8)$	$0.080_7(9_4)$	$0.091_8(7_3)$
14	10	4	15	11	5	$1102.54508_0(7_0)$	$0.083_9(6_2)$	$0.073_7(2_6)$	$0.081_8(5_4)$
18	9	9	19	10	10	$1103.1237_6(1_0)$	$0.0881_9(6_6)$	$0.063_2(6_1)$	$0.082_9(1_8)$
23	7	17	24	8	16	$1106.3885_6(1_2)^d$	$0.086_8(8_9)$	$0.072_5(8_8)$	$0.083_8(8_9)$
18	8	10	19	9	11	$1106.49250_5(7_5)$	$0.088_2(5_8)$	$0.071_3(7_6)$	$0.084_7(6_2)$
13	9	5	14	10	4	$1106.56840_5(2_5)$	$0.088_2(6_6)$	$0.068_0(8_2)$	$0.084_0(6_9)$
19	17	3	18	16	2	1216.82777(13)	$0.090_9(9_3)$	$0.080_2(7_8)$	$0.088_{7}(9_{0})$
33	16	18	32	15	17	$1221.3600_8(4_8)$	$0.099_1(4_6)$	$0.068_7(9_1)$	$0.092_7(5_5)$
30	18	12	29	17	13	$1226.32107_0(8_0)$	$0.101_0(8_1)$	$0.081_2(1_7)$	$0.096_8(6_8)$
$^{34}S^{16}O_2$									
J'	$K_{a}^{'}$	$K_{c}^{'}$	J"	$K_a^{"}$	$K_c^{"}$	$\nu_0{}^b$	$\gamma_{\mathrm{N}_2}^{0}{}^{c}$	$\gamma_{{ m O}_2}^{0}{}^c$	$\gamma_{ m air}^{0-c}$
25	13	13	26	14	12	1083.39323(20)	$0.091_6(1_0)$	0.0848(19)	$0.090_2(1_2)$
16	15	1	17	16	2	$1083.6045_3(2_4)$	$0.098_7(3_0)$	$0.072_1(9_0)$	$0.093_1(4_3)$
29	12	18	30	13	17	$1083.6126_4(2_7)$	$0.118_2(7_6)$	$0.080_6(7_7)$	$0.110_3(7_6)$
25	7	19	26	8	18	$1102.4451_6(1_2)$	$0.092_4(5_8)$	$0.067_0(2_7)$	$0.087_1(5_1)$
15	9	7	16	10	6	1102.9936 <sub>3</sub> (3 <sub>6</sub> )	$0.091_4(4_2)$	$0.0604_9(7_7)$	$0.084_9(3_5)$
25	11	15	25	12	14	$1106.2350_1(7_1)$	$0.109_7(3_6)$		

<sup>&</sup>lt;sup>a</sup> Uncertainties in parentheses are one standard deviation.

<sup>b</sup> Transition frequency in cm<sup>-1</sup>.

<sup>c</sup> Broadening coefficient in cm<sup>-1</sup> atm<sup>-1</sup>.

<sup>&</sup>lt;sup>d</sup> Overlapped to  $42_{11,31} \leftarrow 42_{12,30} v_1 + v_2 - v_2^{32} S^{16} O_2$ .

Table 3. Assignments, transition frequencies,  $N_2$ - and  $O_2$ -broadening coefficients of pure rotational lines of  $^{32}S^{16}O_2$ .

J'	$K_{a}^{'}$	$K_{c}^{'}$	J"	$K_a^{"}$	$K_c$	$v_0{}^b$	$\gamma_{\mathrm{N}_2}^{0}$	$\gamma_{\mathrm{O}_2}^{0}{}^{c}$	$\gamma_{ m air}^{0}$	
J dependence										
11	11	1	11	10	2	1064050.059(50)	$0.096_6(1_3)$	$0.078_3(1_5)$	$0.0927(1_3)$	
12	11	1	12	10	2	1064089.718(50)	$0.0986_{7}(7_{4})$	$0.081_0(1_6)$	$0.0949_5(9_3)$	
15	11	5	15	10	6	$1064206.09_6(5_0)$	$0.1038_6(7_1)$	$0.084_3(1_2)$	$0.0997_5(8_1)$	
16	11	5	16	10	6	1064241.229(50)	$0.103_7(1_1)$	$0.083_9(1_1)$	$0.099_6(1_1)$	
26	11	15	26	10	16	$1064223.09_5(5_0)$	$0.1011_2(7_4)$	$0.07557(6_1)$	$0.0957_6(7_1)$	
28	11	17	28	10	18	$1064072.08_0(5_0)$	$0.096_2(1_3)$	$0.0758_8(8_4)$	$0.091_9(1_2)$	
34	11	23	34	10	24	1063097.643(50)	$0.099_2(1_2)$	$0.0729_6(5_8)$	$0.0937(1_1)$	
42	11	31	42	10	32	1059981.907(50)	$0.097_9(1_7)$	$0.070_8(1_2)$	$0.092_2(1_6)$	
51	11	41	51	10	42	1052515.119(50)	$0.0991_2(6_8)$	$0.0680_9(7_1)$	$0.0926_{1}(6_{9})$	
60	11	49	60	10	50	1038159.319(50)	$0.0954_2(7_1)$	$0.0631_5(7_6)$	$0.0886_4(7_2)$	
$K_a$ d	lepend	lence								
10	1	9	10	0	10	$104239.29_5(2_0)$	$0.1151_4(3_0)$	$0.0789_4(7_1)$	$0.1075_4(3_9)$	
10	2	8	10	1	9	$129514.80_1(2_0)$	$0.1143_1(7_6)$	$0.0796_8(2_8)$	$0.1070_3(6_6)$	
10	4	6	10	3	7	356755.187(20)	$0.1106_8(2_8)$	$0.0814_3(3_8)$	$0.1045_4(3_0)$	
10	5	5	10	4	6	459668.356(20)	$0.1089_3(3_8)$	$0.0826_9(3_5)$	$0.1034_2(3_7)$	
10	6	4	10	5	5	561560.31 <sub>3</sub> (2 <sub>0</sub> )	$0.1071_1(7_6)$	$0.0870_3(7_4)$	$0.1028_9(7_6)$	
10	7	3	10	6	4	$662997.74_9(2_0)$	$0.1041_2(7_9)$	$0.0920_5(7_9)$	$0.1015_8(7_9)$	
$\Delta J$ =	= -1									
20	7	13	21	6	16	$258666.957(2_0)$	$0.0951_7(8_6)$	$0.0743_5(3_0)$	$0.0908_0(74)$	
$\Delta J$ =	= +1									
43	5	39	42	6	36	248995.12 <sub>1</sub> (2 <sub>0</sub> )	$0.0967_4(2_0)$	$0.0670_3(3_8)$	$0.0905_0(2_4)$	
3	3	1	2	2	0	$313279.72_9(2_0)$	$0.1121_5(2_5)$	$0.0818_1(3_0)$	$0.1057_8(2_6)$	

 <sup>&</sup>lt;sup>a</sup> Uncertainties in parentheses are three times the standard deviation.
 <sup>b</sup> Transition frequencies in MHz.
 <sup>c</sup> Broadening coefficient in cm<sup>-1</sup> atm<sup>-1</sup>.